

Pushing the Limits of K2: Observing Trans-Neptunian Objects

S3K2: Solar System Studies with K2

Róbert Szabó¹

collaborators:

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Goals and methods

- Group of Hungarian senior and young astronomers, experts in **Kepler/K2 photometry**, variable stars, exoplanets, **infrared photometry**, **comets/asteroids**

WHO

- **Solar System** studies with Kepler/K2
- pushing the limits:
 - **detection of faint and moving objects**
Trans-Neptunian objects + Main-Belt Asteroids + moons
- study rotation rate, surface variegations, companions

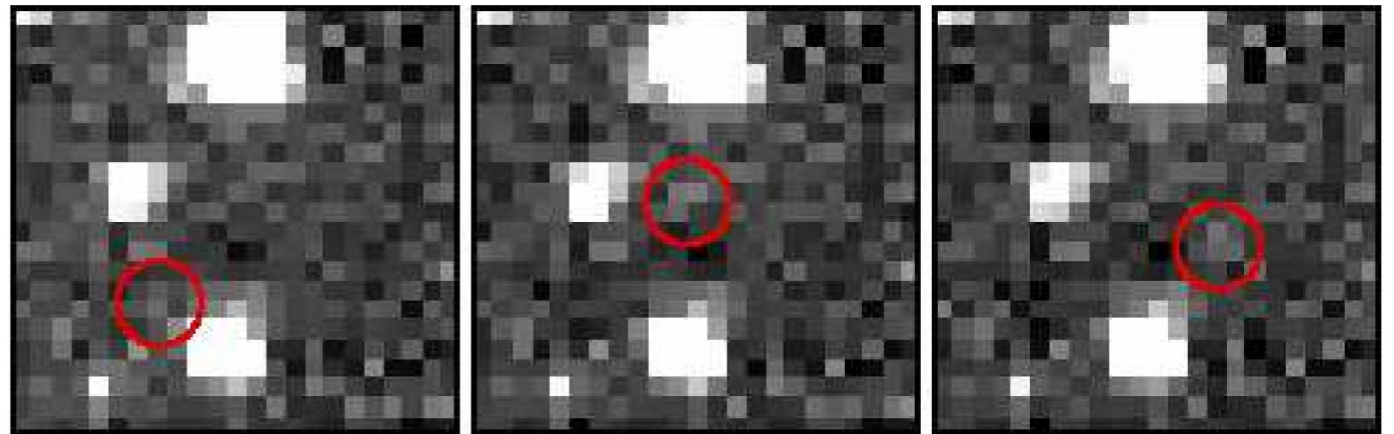
WHAT

- **boomerang-shaped pixel masks**
around the stationary point (TNOs)
- large, contiguous **supermasks** (open clusters, etc.)
- astrometry + optical photometry:
 - **fitsh** package (Pál 2012, <http://fitsh.szofi.net>)
- **infrared**: Herschel “TNOs are Cool!” key program
- **thermophysical modeling**
thermal flux + abs. magnitude + rotation —>
diameter + geom. albedo
- **White Paper Kiss Cs. et al. 2013**

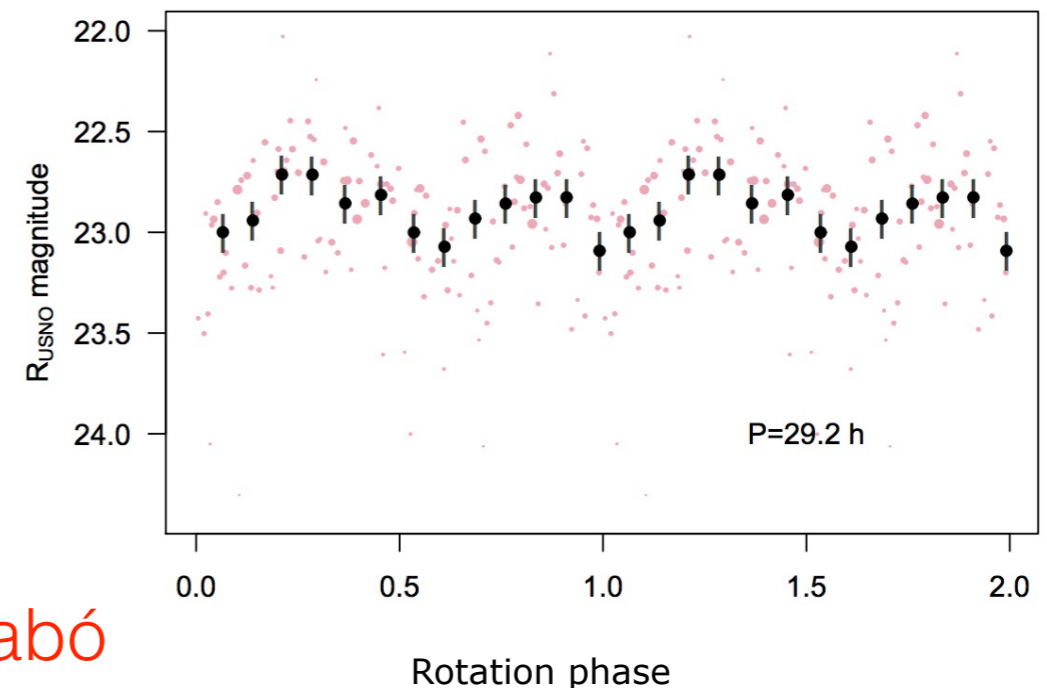
HOW

TNO 2002 GV₃₁ in C1

- classical Kuiper-belt object
- 40 - 48 AU $P_{\text{orb}} = 295\text{yr}$
- $R = \mathbf{22.5 - 23.0}$ mag
- Herschel: no detection
- $d < 180$ km
- K2: averaged 5 measurements: resulting in **129** data points
- **16** days around the stationary point
- **Detection!**
- First light curve information
- $P_{\text{rot}} = \mathbf{29.2 \pm 1.1}$ h
- long rotation period
most TNOs have $P_{\text{rot}} < 1$ day



1-day co-added images

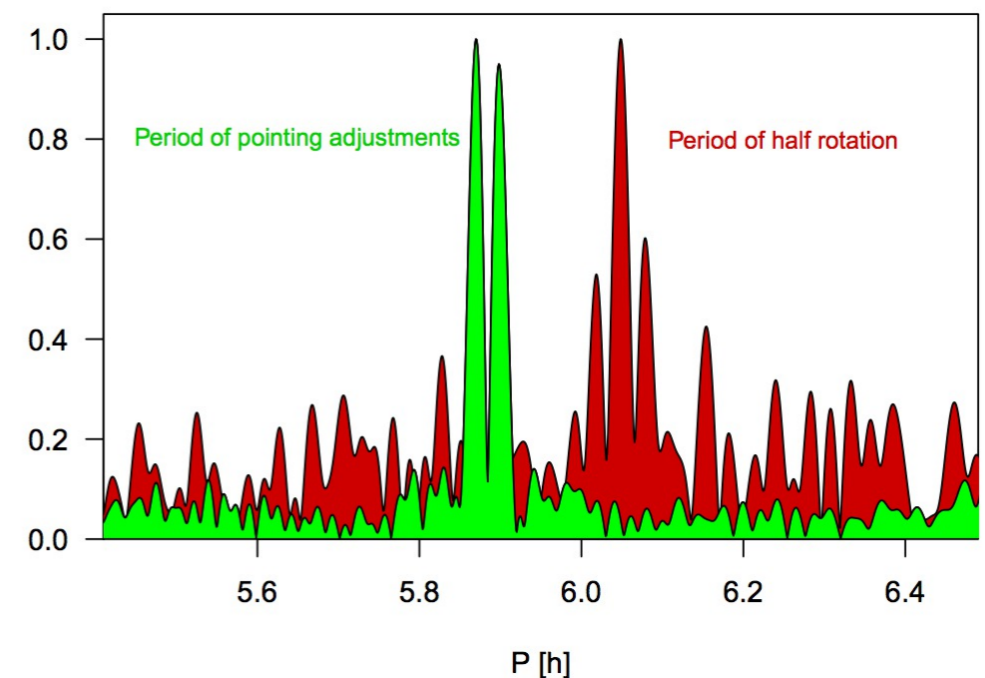
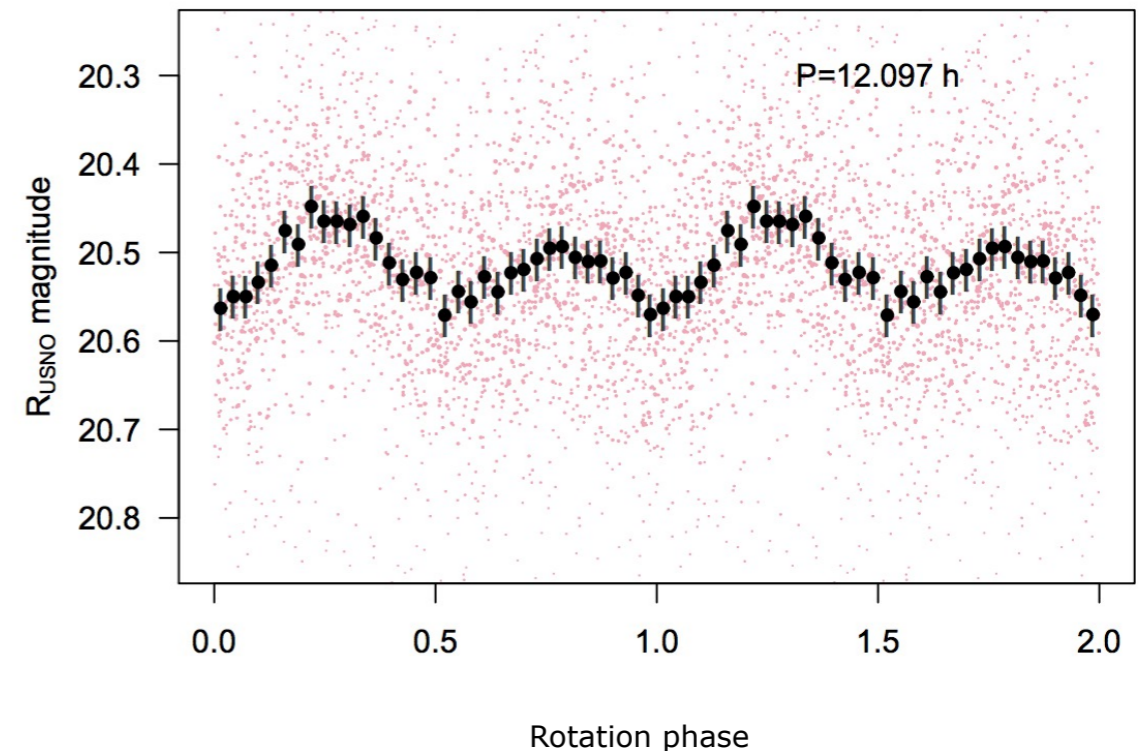


White Paper Kiss Cs. et al., GO1064 Szabó
Pál et al. ApJ, 804, L45, 2015

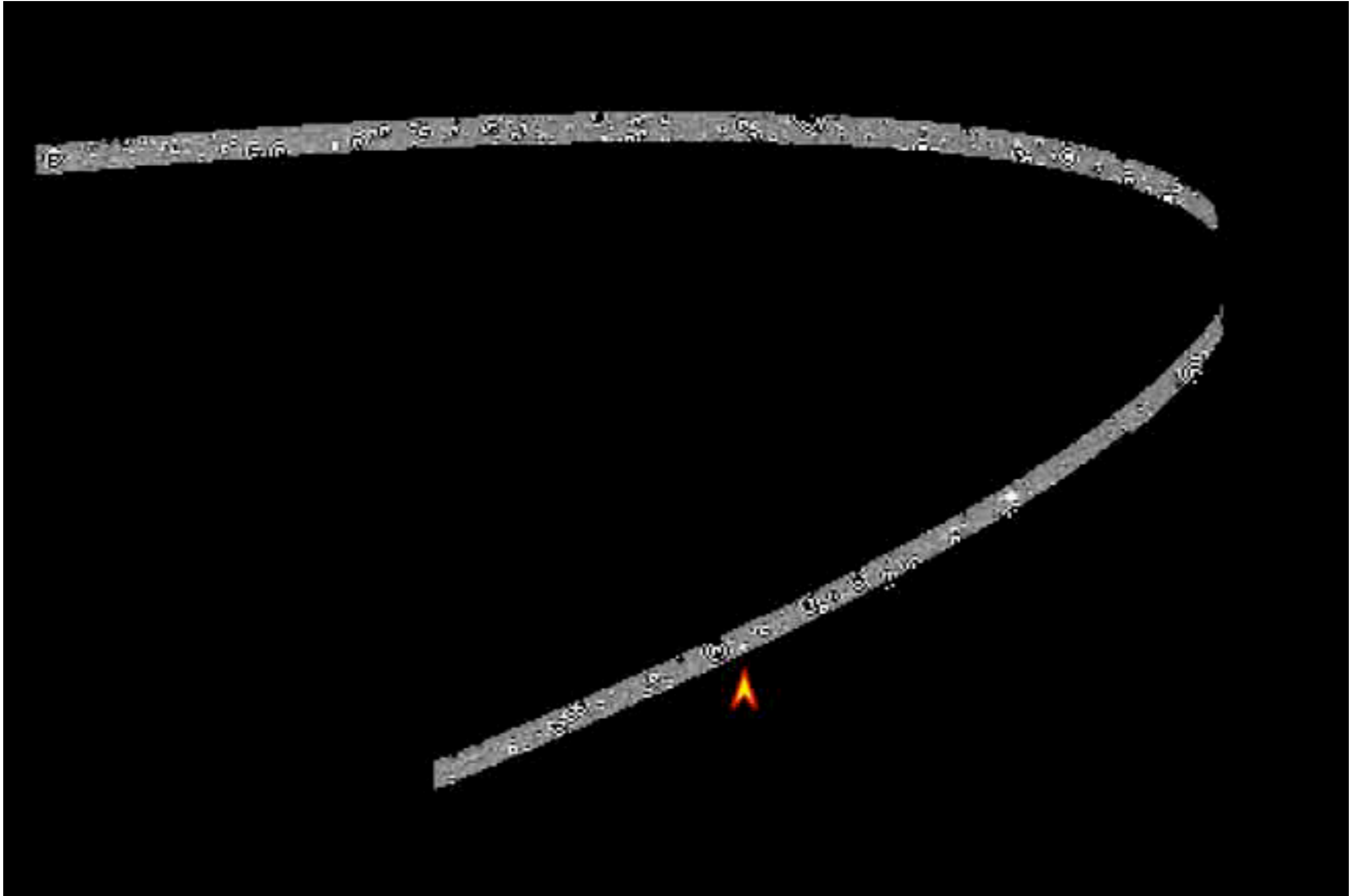
(278361) 2007 JJ₄₃ in C2

- outer Kuiper-belt object
- 40-56 AU $e=0.159$ $i=12.0^\circ$
- likely in 2:1 resonance with Neptune
- $d = 670 \pm 150$ km
- asymmetric phased light curve
- $P_{\text{rot}} = \mathbf{12.097 \text{ h}}$
- rot. period clearly separated from the pointing adjustments
- Benecchi & Sheppard (2013) rot. period confirmed
- uneven surface structures / shape

GO2066 Schwamb, GO3053 Szabó
Pál et al. ApJ, 804, L45, 2015



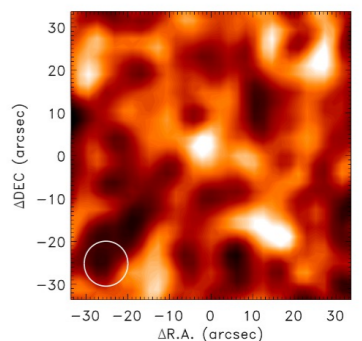
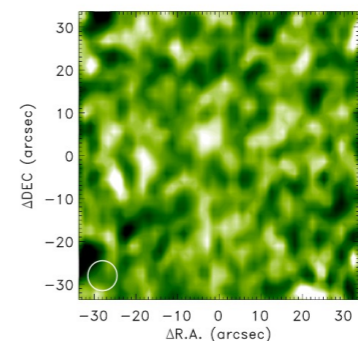
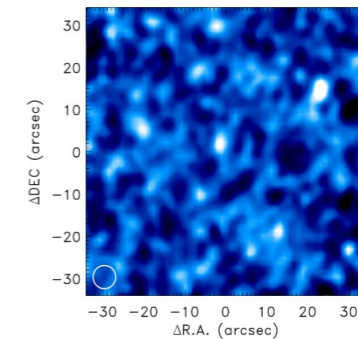
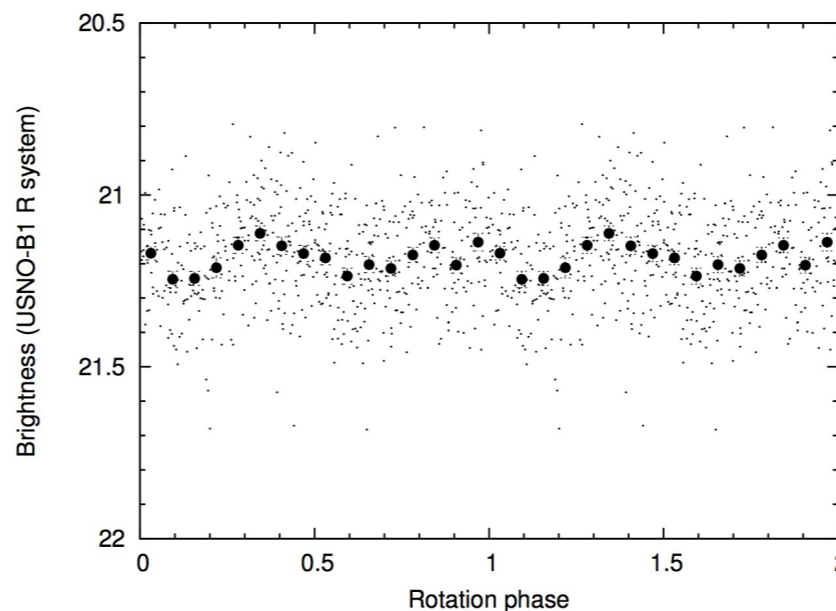
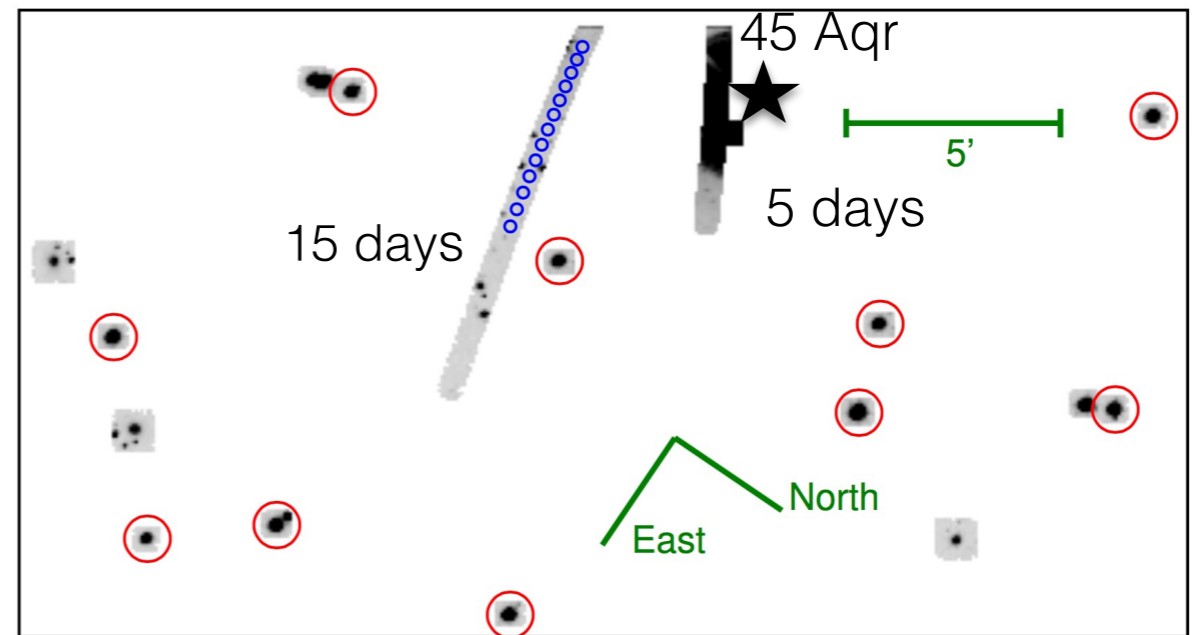
2007JJ₄₃



(225088) 2007 OR₁₀ in C3

- second most distant TNO
87 AU, 33-108 AU
- scattered disk object
- **first optical rotational variation** detected with K2
- $P_{\text{rot}} = \mathbf{44.81 \text{ h}}$ greatly affects heat distribution
- Herschel 60-210 μm \rightarrow
alb. $p_v = 0.089^{+0.031}_{-0.009}$
size $d = \mathbf{1535}^{+75}_{-225} \text{ km}$
- larger, darker
- **3rd largest dwarf planet!**
- volatiles (CH₄, CO, N₂) may be retained

GO3053 Szabó R. et al.
Pál et al. AJ, submitted

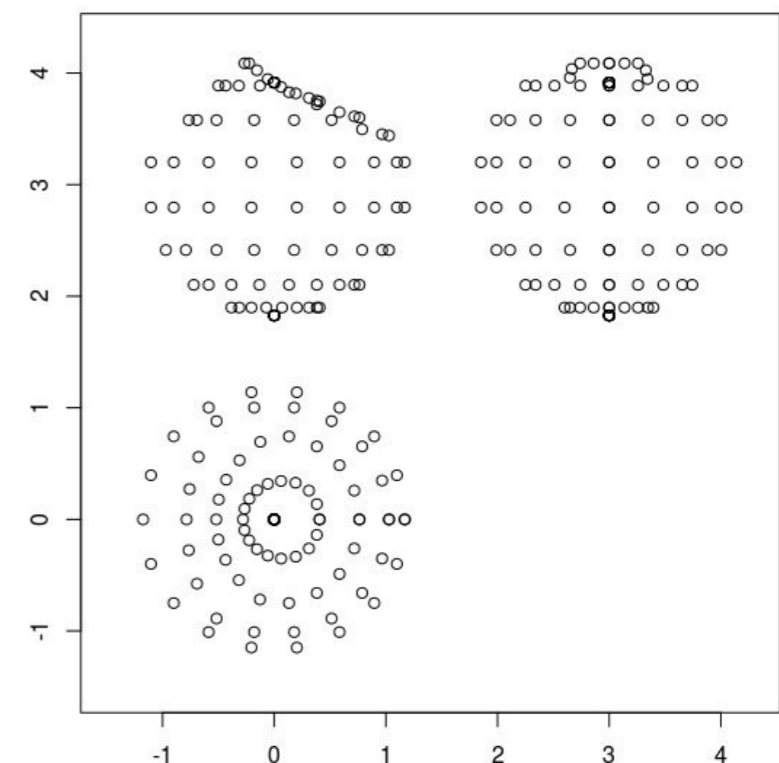
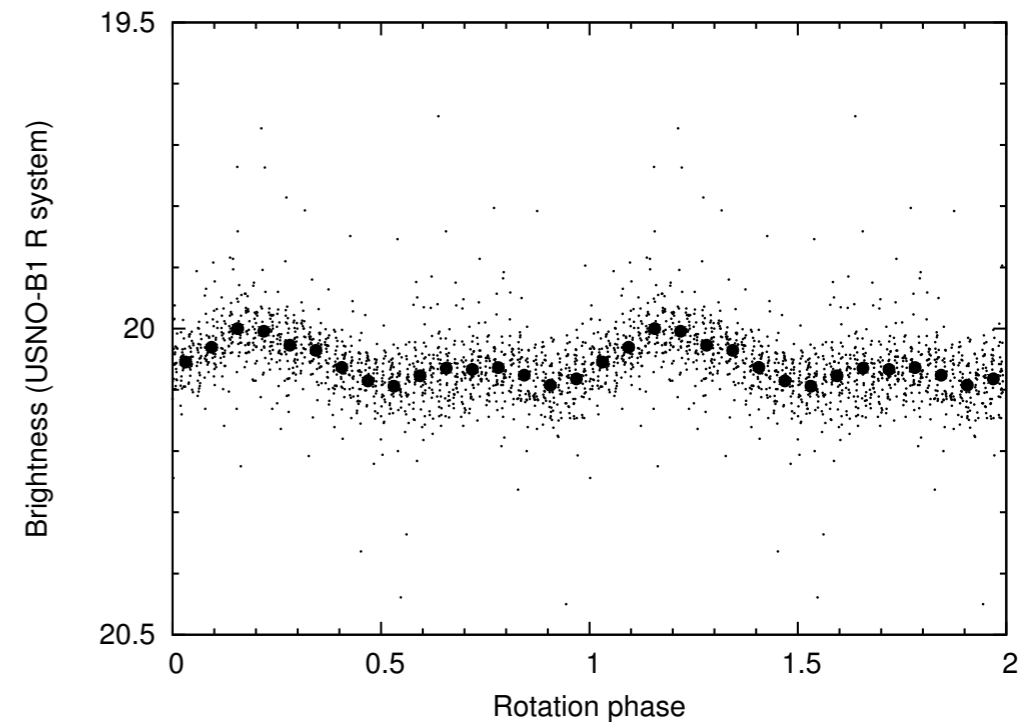


(250112) 2002 KY₁₄ in C4

- observed for **29** days
- **20th** magnitude
- distance between 8.6-16.5 AU
- $d=47\pm 4$ km Duffard et al. 2014

- $P_{\text{rot}} = \mathbf{8.502\ h}$
- double-peaked maxima
- asymmetric light curve
needs a planar cut in the shape
- **first centaur with a 3D shape model**

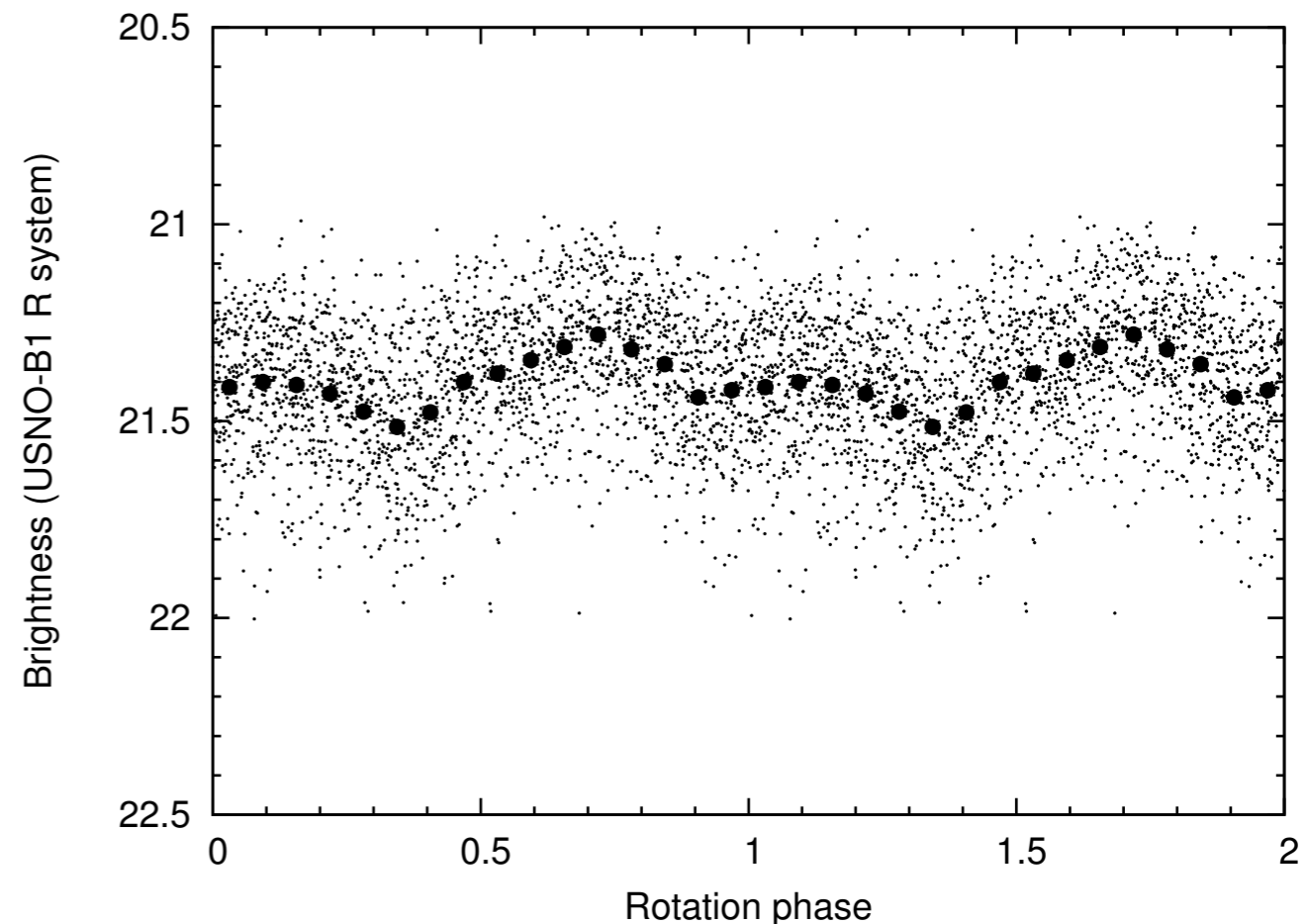
GO04110 Schwamb
Szabó Gy. M. et al. in prep.



(126154) 2001 YH₁₄₀ in C5

- observed for **77** days
- **21.3** magnitude
- distance between 36-48 AU
- 3:5 resonance with Neptune
d= 300-390 km
- $P_{\text{rot}} = \mathbf{27.38 \pm 0.14 \text{ h}}$
- double-peaked maxima
at >15 sigma confidence level
- Good agreement with
Sheppard 2007 AJ, **134**, 787
($P_{\text{rot}} = \mathbf{13.25 \pm 0.2 \text{ h}}$)

GO05110 Schwamb
work in progress !



Nereid, moon of Neptune in C3

- large, irregular satellite
- captured TNO?
- very eccentric and irregular orbit
- shape, orientation, rotation?
- small / large amplitude phases
- $P_{\text{rot}} = 11.504 \pm 0.017 \text{ h}$
- max. a:c axis ratio: **1.3 : 1.0**
cannot be in forced precession state
- low amplitude state
- Herschel + Spitzer observations
- **high surface roughness** (cratered)
- $d = 345 \pm 15 \text{ km}$

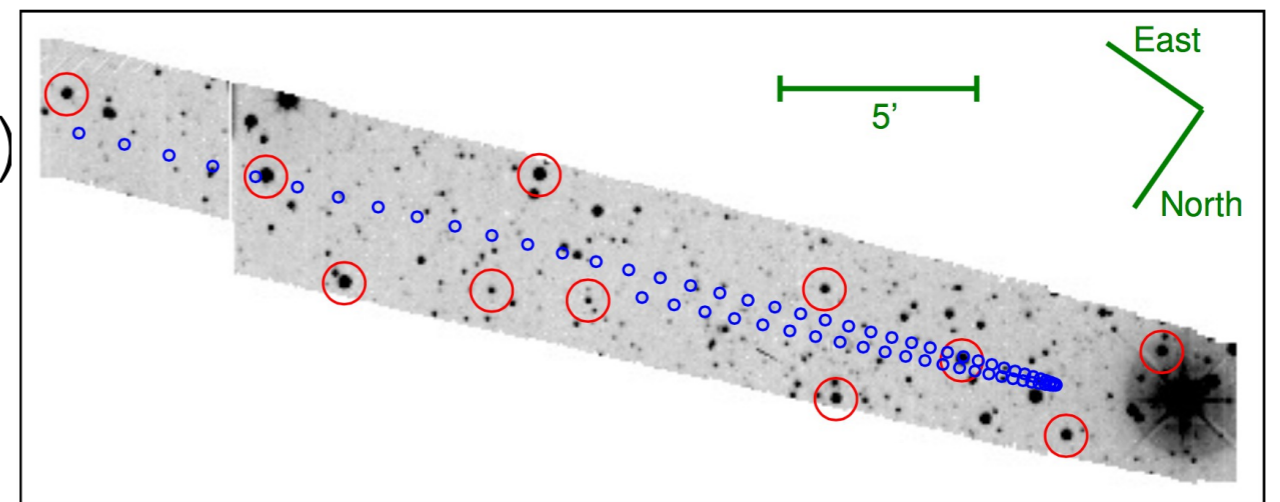
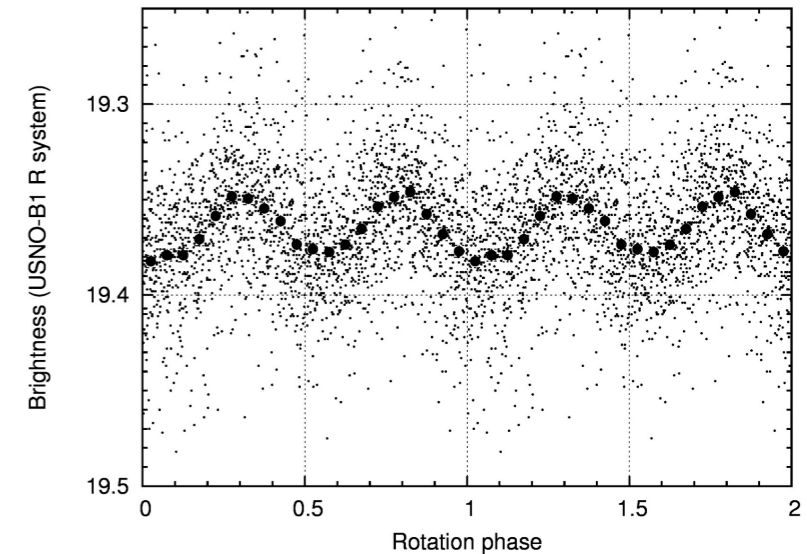
GO3115 Schaefer

Kiss Cs. et al. MNRAS, submitted

Best Voyager-2 image 1989



K2 light curve



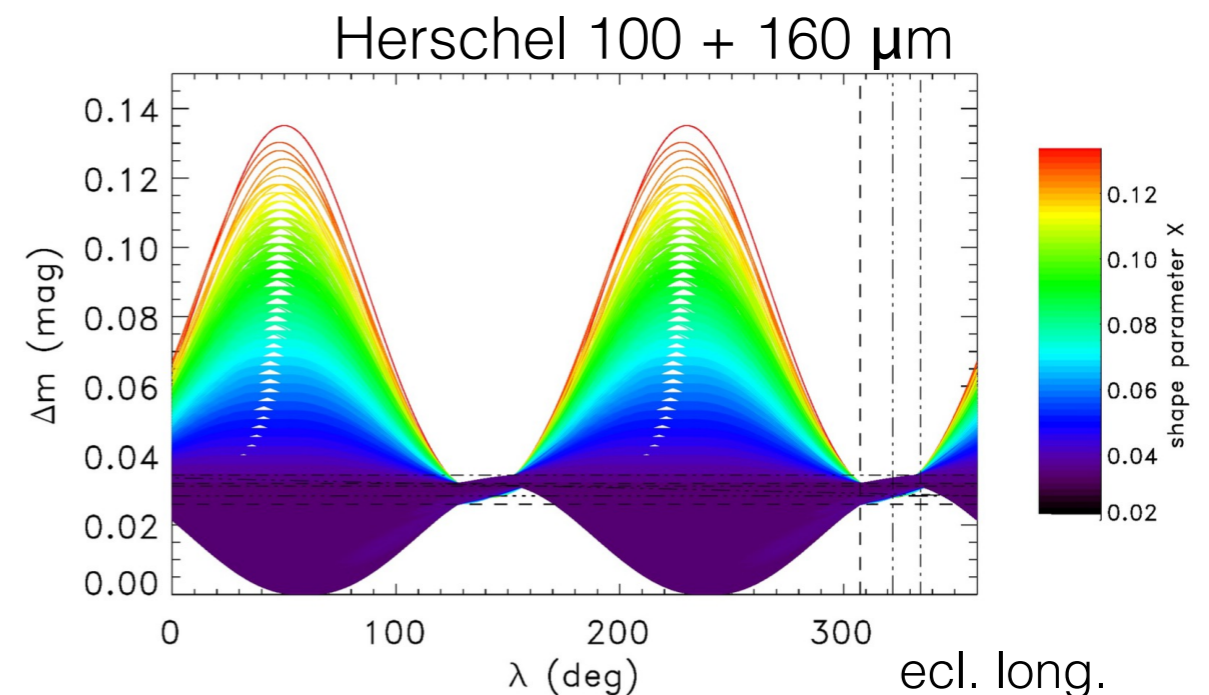
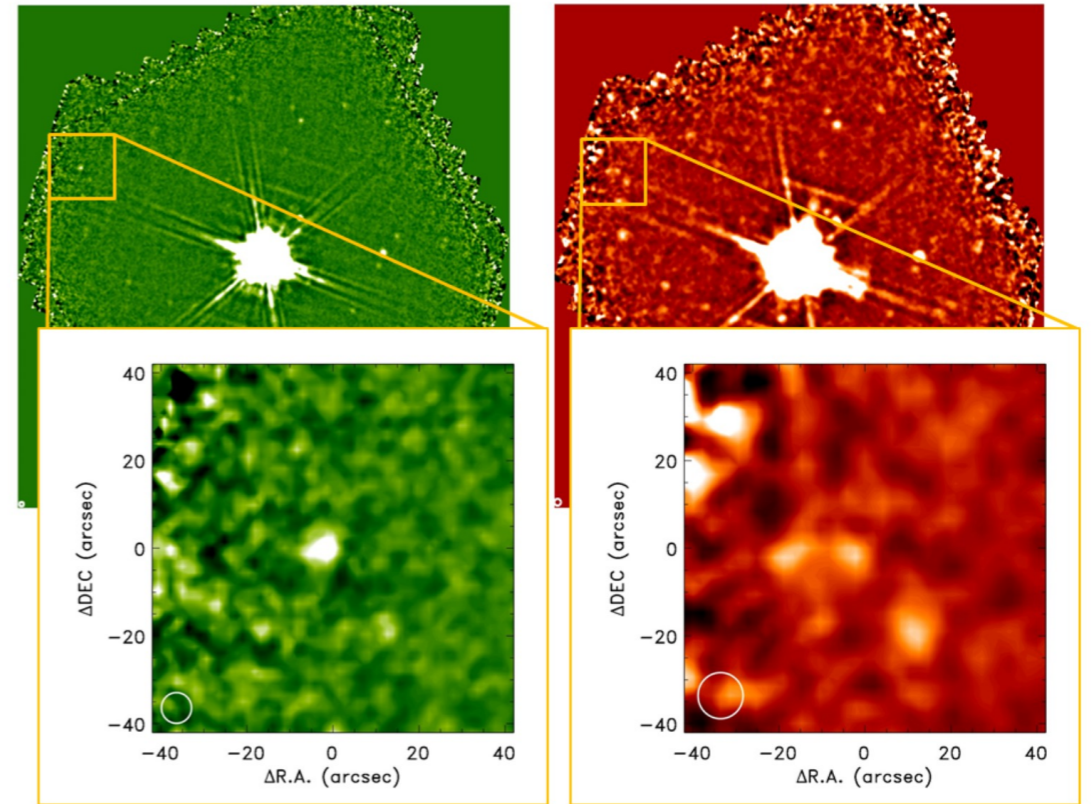
Path of Neptune and Nereid in the K2 FoV

Nereid, moon of Neptune in C3

- large, irregular satellite
- captured TNO?
- very eccentric and irregular orbit
- shape, orientation, rotation?
- small / large amplitude phases
- $P_{\text{rot}} = 11.504 \pm 0.017 \text{ h}$
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GO3115 Schaefer

Kiss Cs. et al. MNRAS, submitted



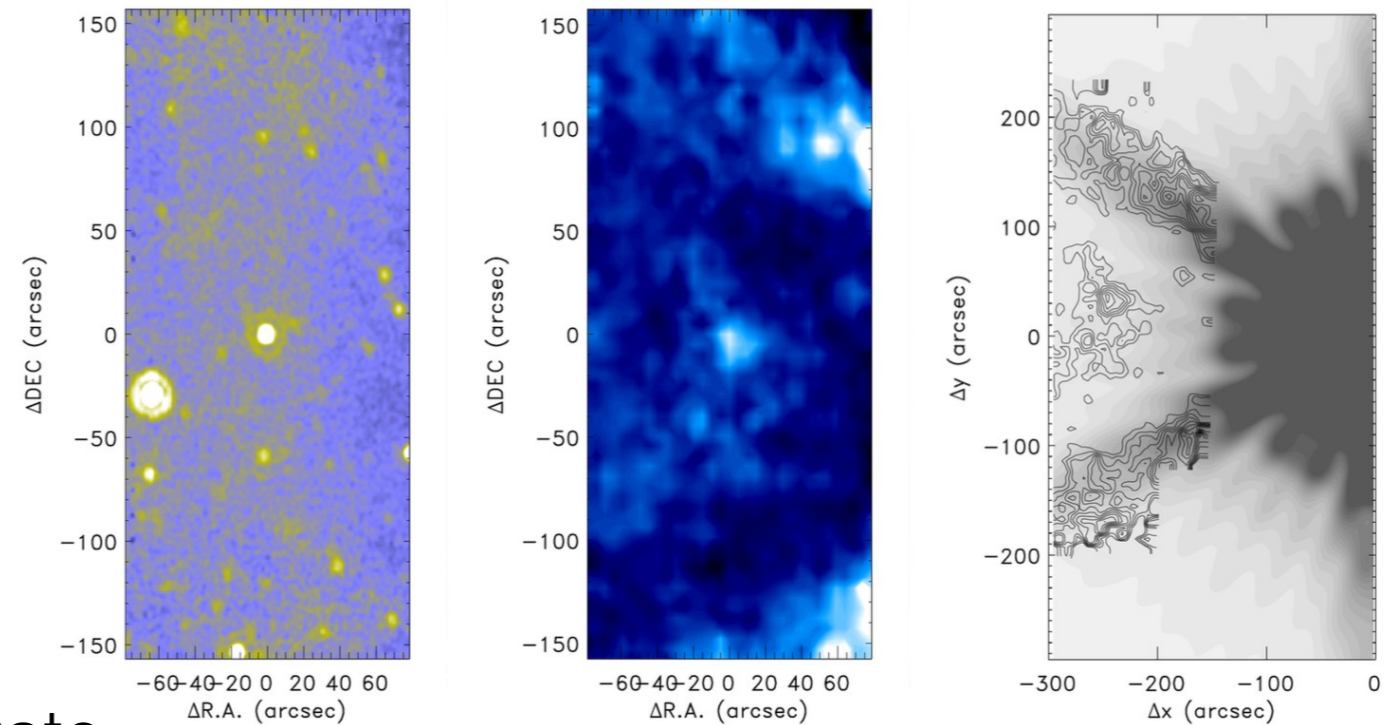
X: triaxial ellipsoid shape parameter $b=1$ $a=(1+X)b$ $c=(1-X)b$

Nereid, moon of Neptune in C3

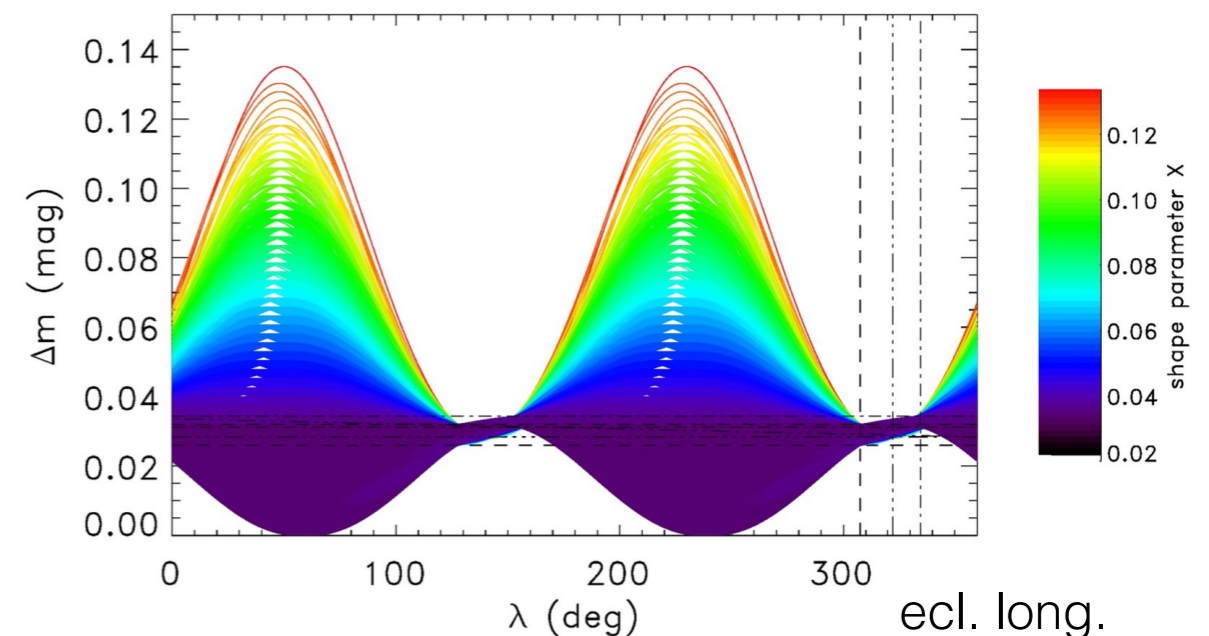
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- captured TNO?
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- max. a:c axis ratio: $\mathbf{1.3 : 1.0}$
cannot be in forced precession state
- low amplitude state
- Herschel + Spitzer observations
- **high surface roughness** (cratered)
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GO3115 Schaefer

Kiss Cs. et al. MNRAS, submitted



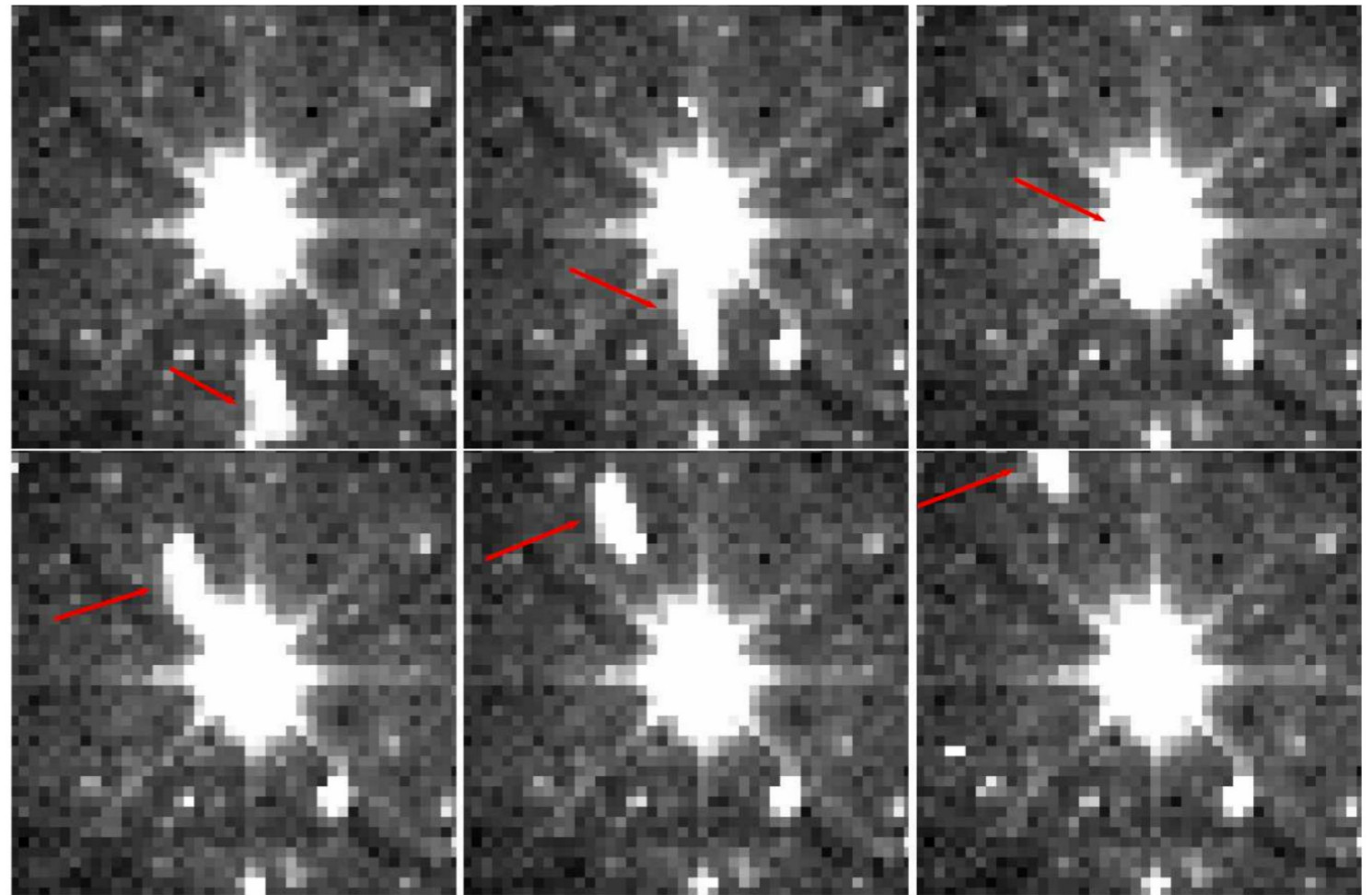
Spitzer 24 + 70 μm



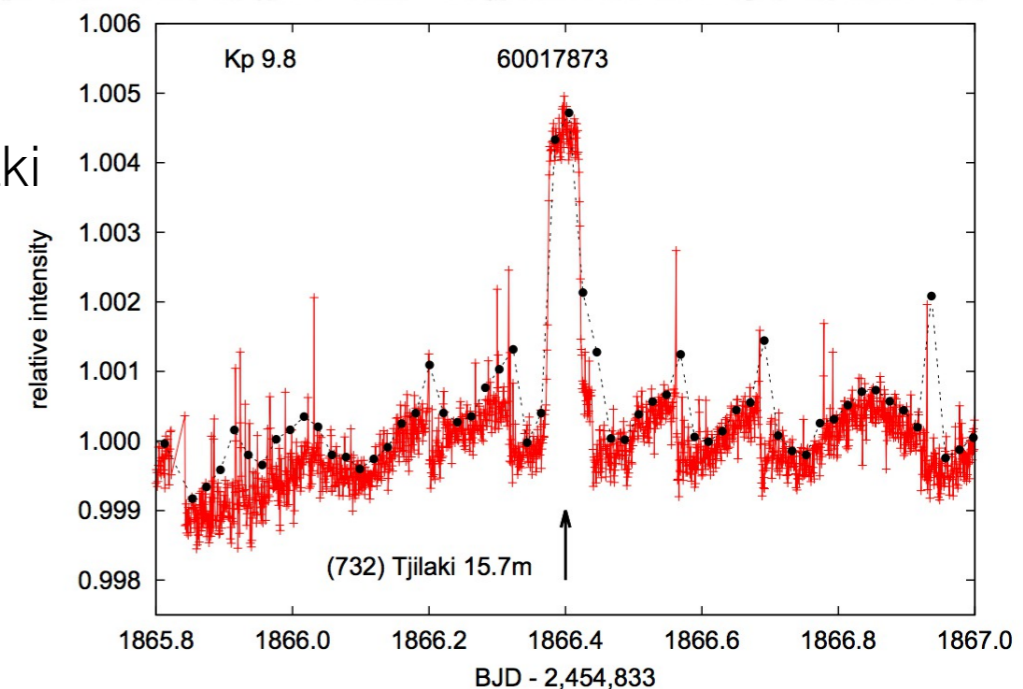
X: triaxial ellipsoid shape parameter $b=1$ $a=(1+X)b$ $c=(1-X)b$

Main-Belt Asteroids in the K2-E2 field

- **2096** targets **8.9** d long
- **300** subsample selected
- **What is the effect of asteroid encounters?**
- All predicted asteroids recovered down to $\sim 21^m$
- All found asteroids could be identified
- No new M-B asteroids



15.7 mag (732) Tjilaki
EPIC 60017873
(Kp=9.8 mag)

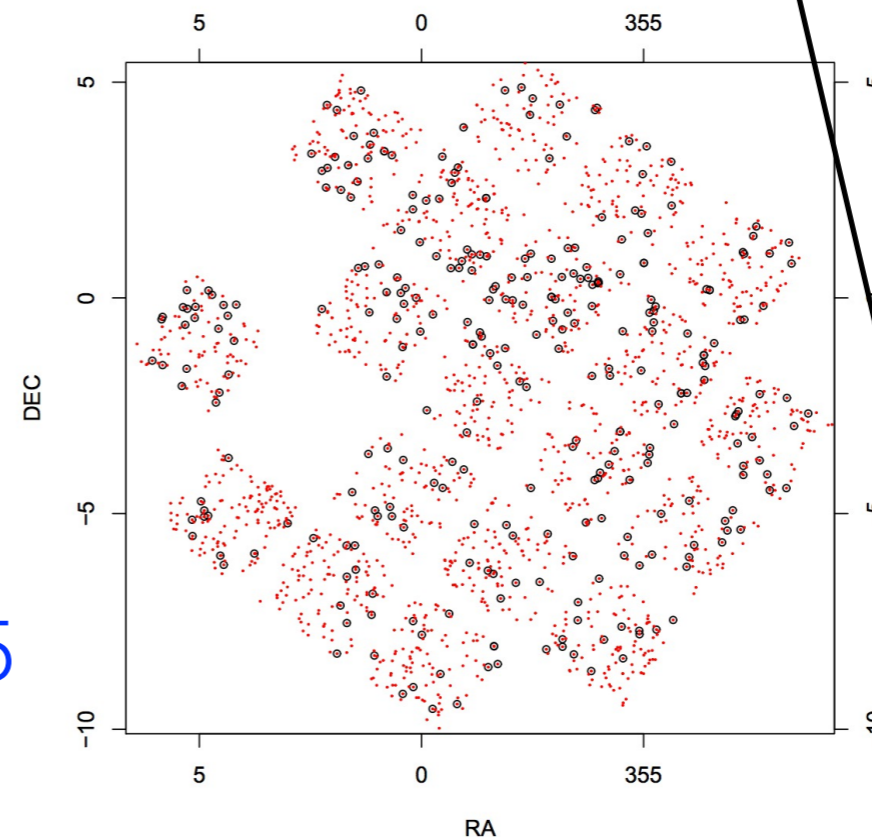
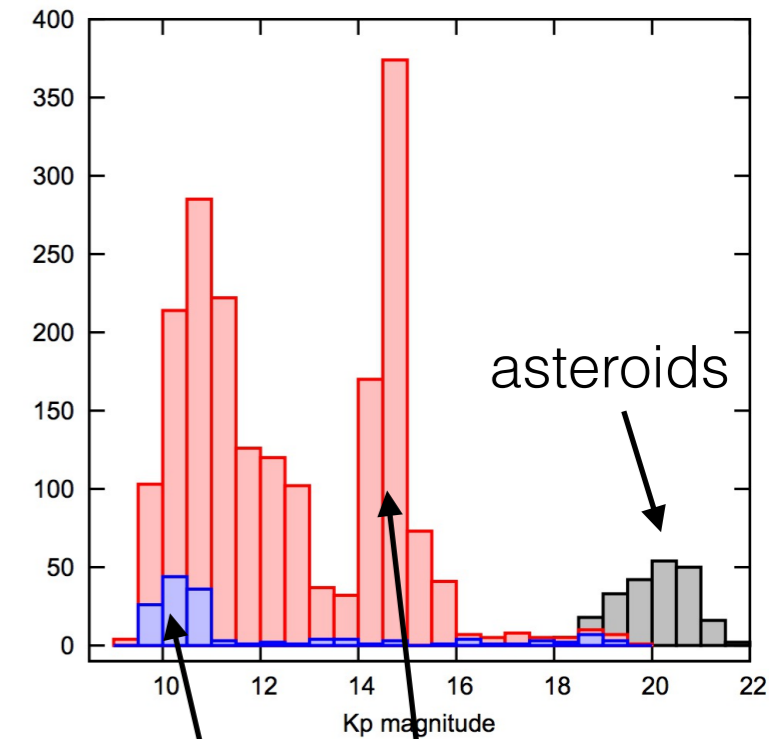
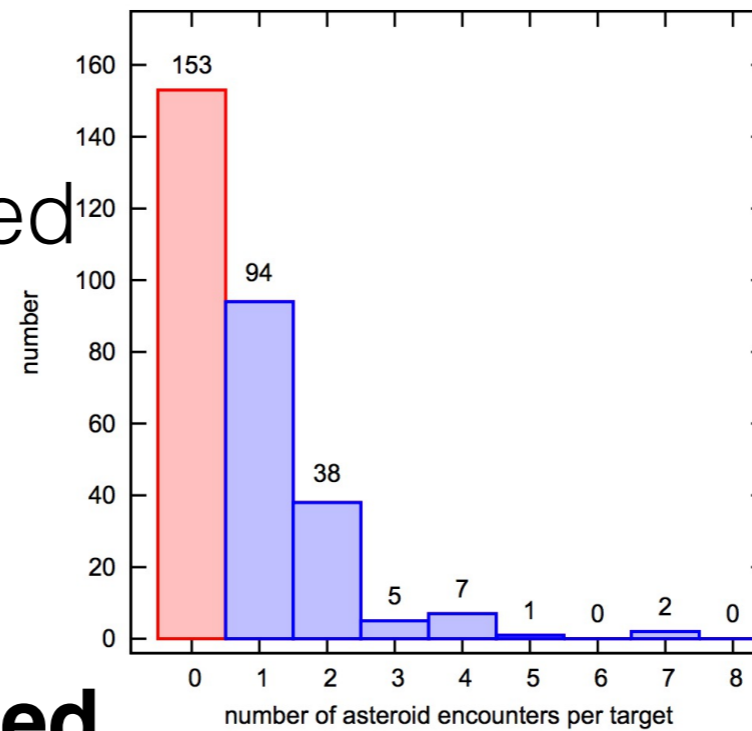


K2 Two-Wheel Concept
Engineering Test

Szabó R. et al. AJ, 149, 112, 2015

Main-Belt Asteroids in the K2-E2 field

- **2096** targets **8.9** d long
- subsample of **300** selected
- **What is the effect of asteroid encounters?**
- **half of our targets showed asteroid encounter!**
- **many targets had several encounters**



red dots:
all K2E2 targets

black circles:
random sample

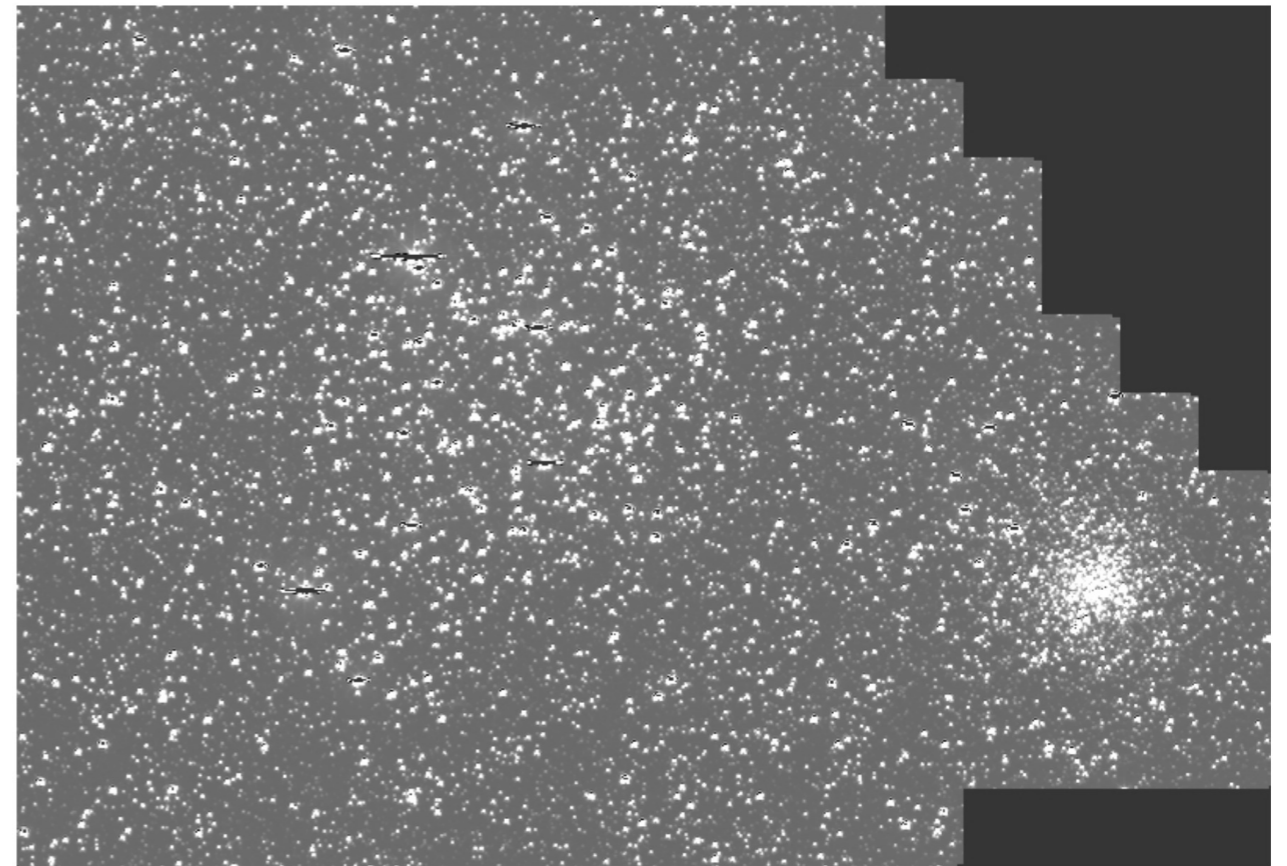
K2 Two-Wheel Concept
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Szabó R. et al. AJ, 149, 112, 2015

K2⁵ Main-Belt Asteroids in the M35 field - C0

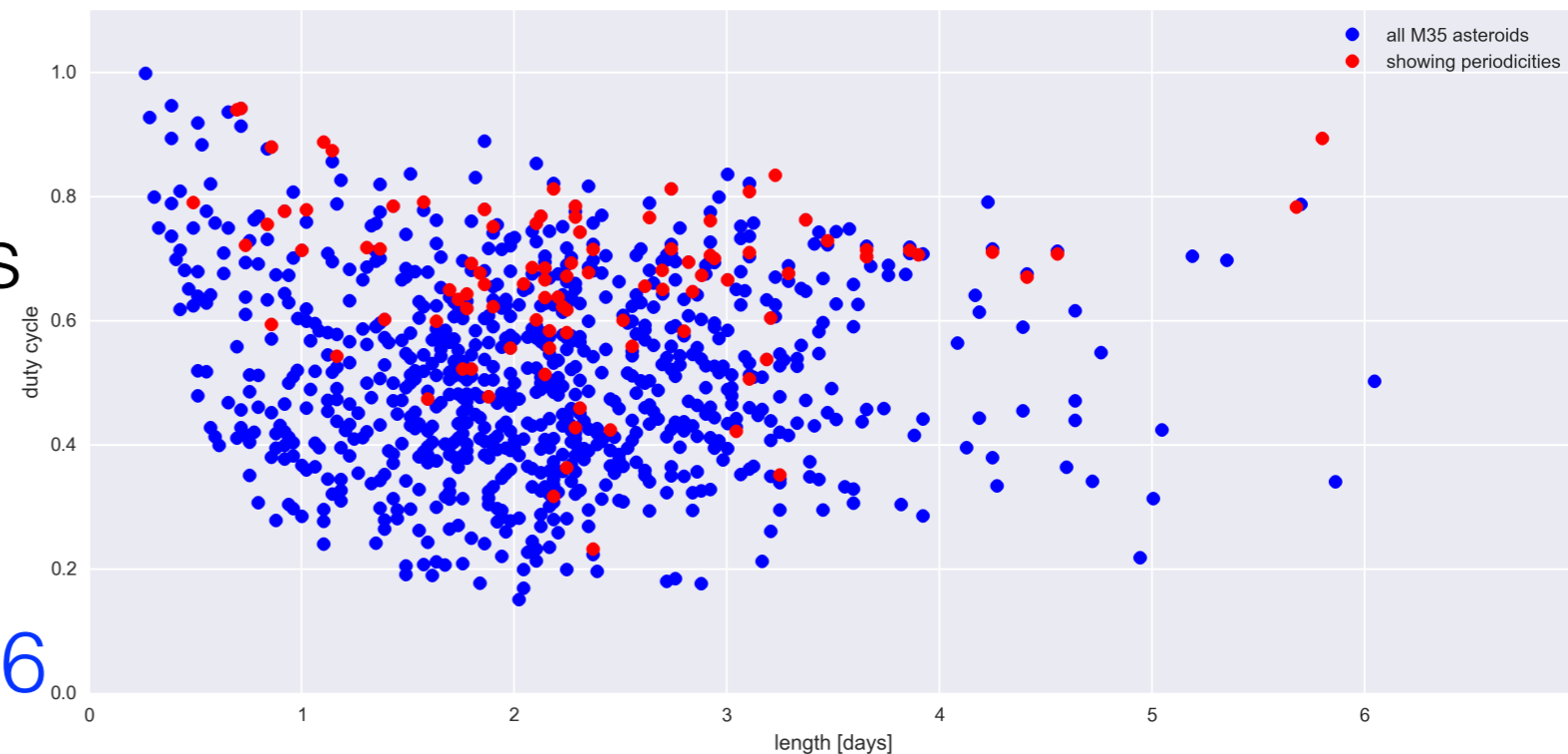
- **924** asteroids crossed the field
- **865** had >12 data points
- **32** showed periodicities
- we used 35-day, continuous fine-point measurements

- Goal: light curves →
shape reconstruction
- $P_{\text{rot}} = \mathbf{2.54 - 56.01}$ h,
median: **7.0** h
- coverage: **0.26 - 6.05** days
median: **2.0** days



M35 superstamp

Szabó R. et al. in prep. 2016

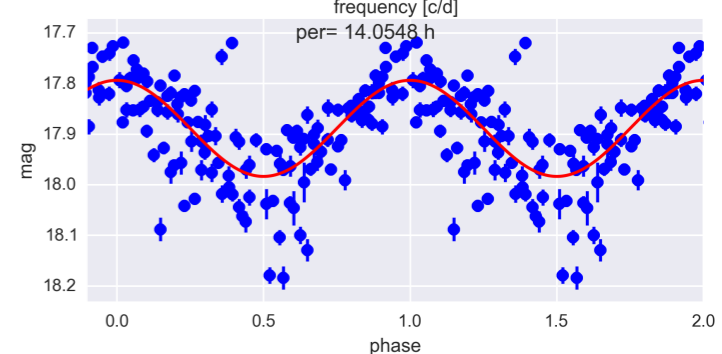
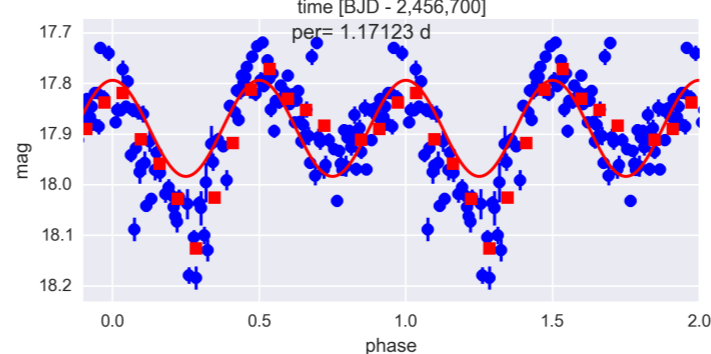
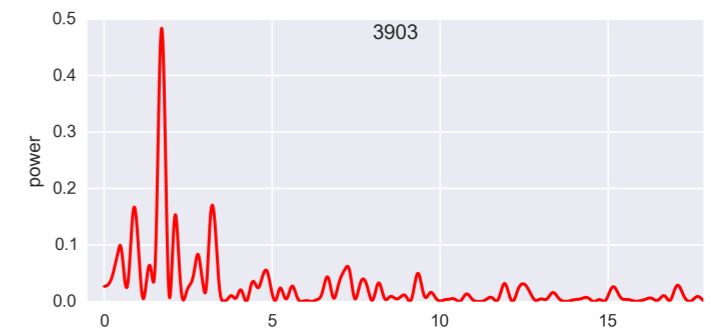
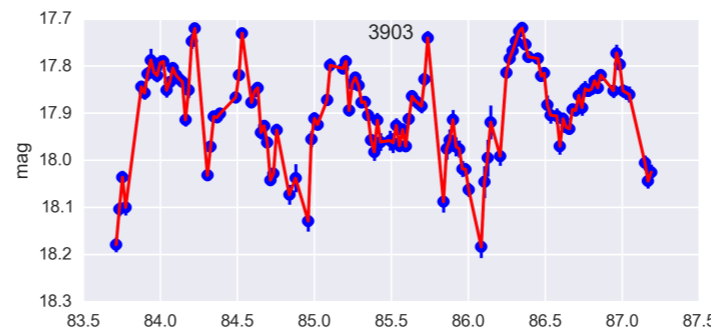
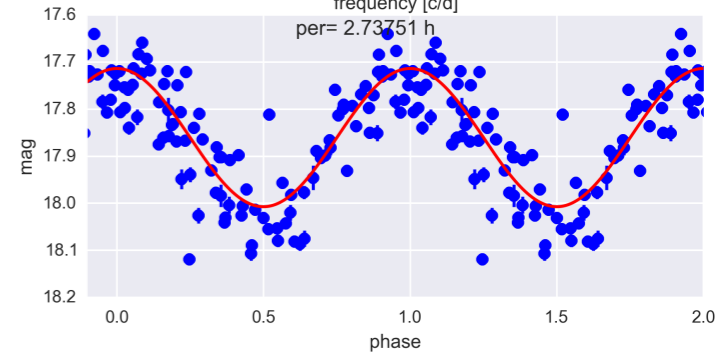
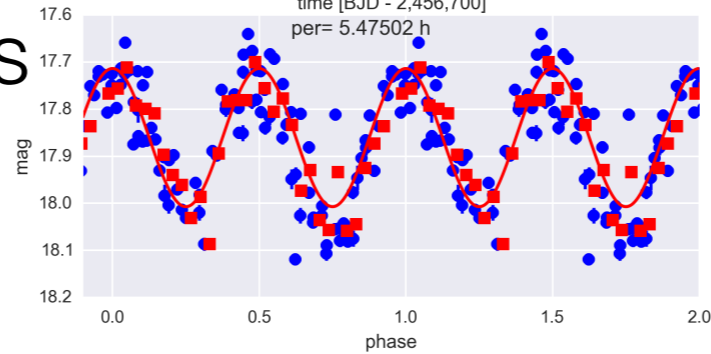
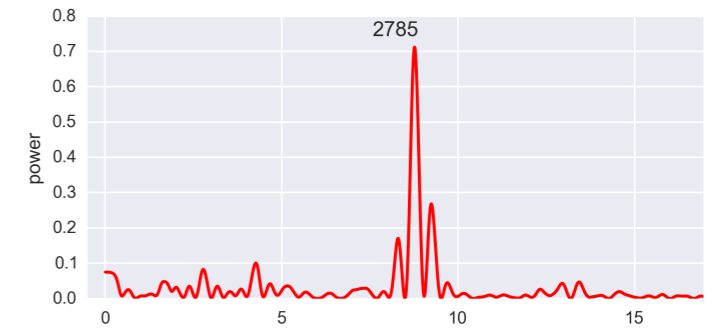
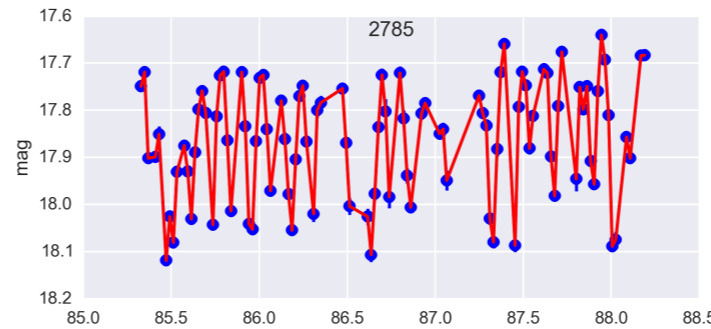


K2⁵ Main-Belt Asteroids in the M35 field - C0

- **924** asteroids crossed the field
- **865** had >12 data points
- **32** showed periodicities
- we used 35-day, continuous fine-point measurements
- Goal: light curves \longrightarrow **shape reconstruction**
- $P_{\text{rot}} = \mathbf{2.54 - 56.01}$ h, median: **7.0** h
- coverage: **0.26 - 6.05** days, median: **2.0** days

M35 superstamp

Szabó R. et al. in prep. 2016



Main-Belt Asteroids in the Nereid field - C3

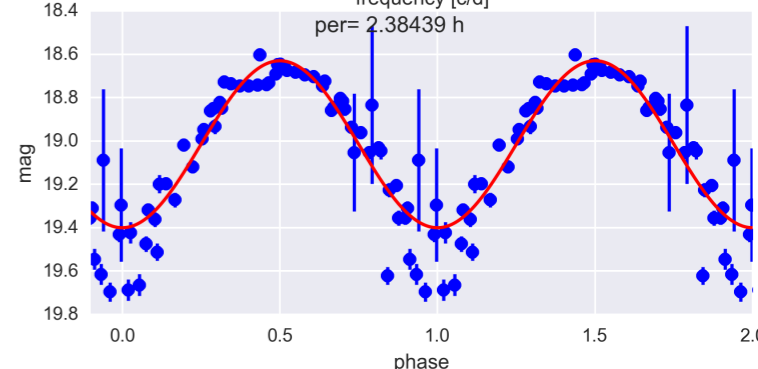
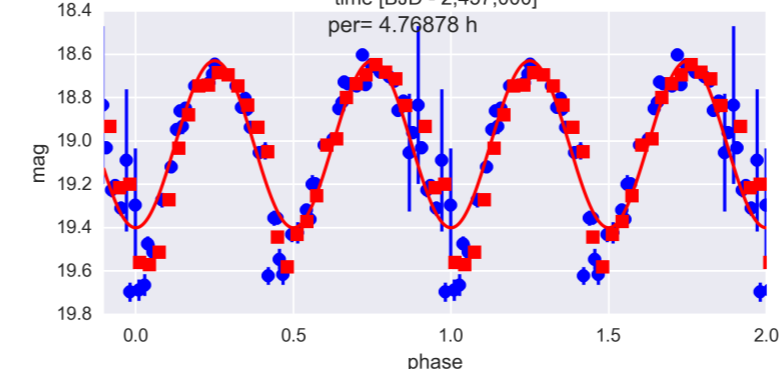
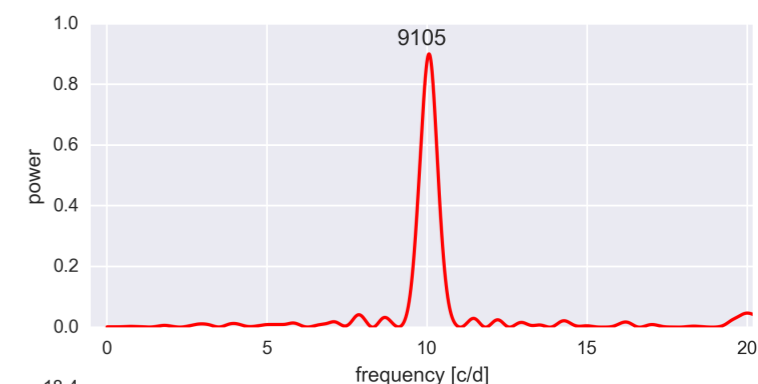
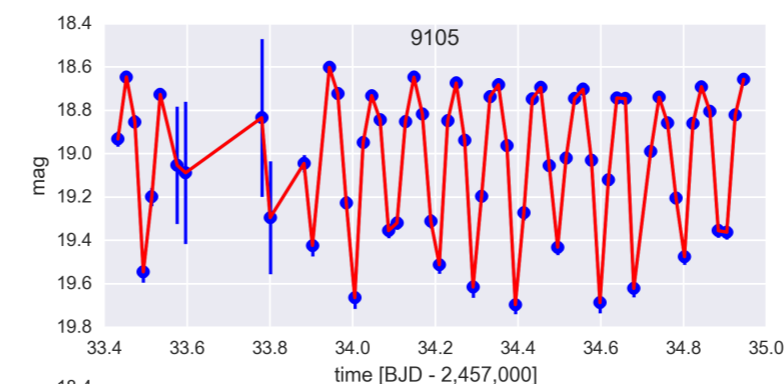
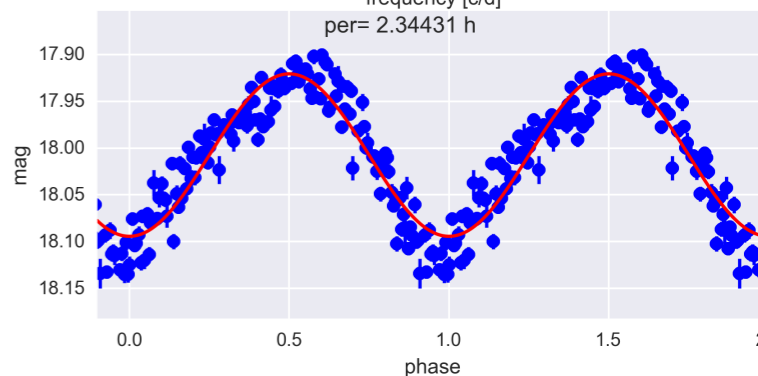
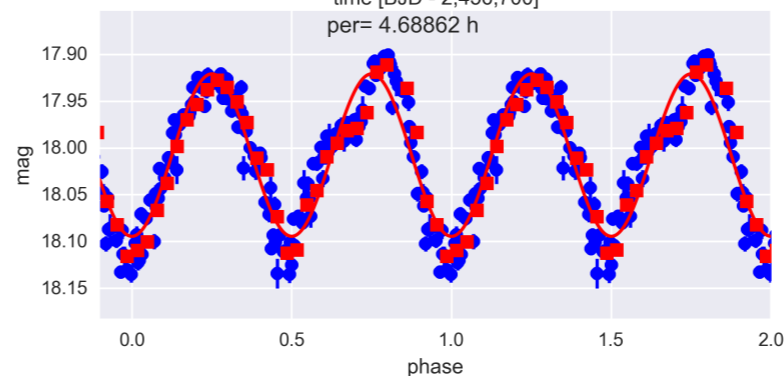
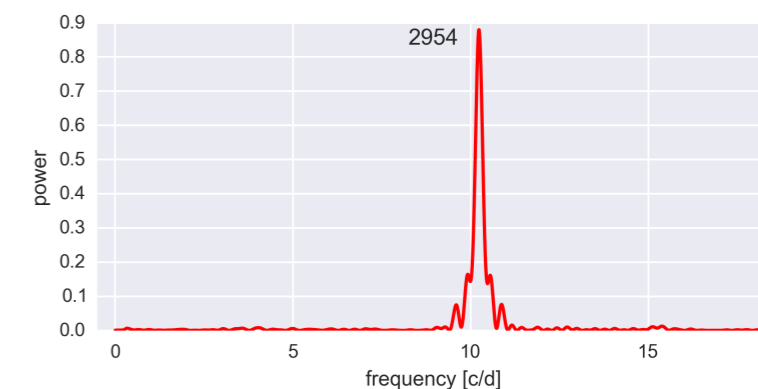
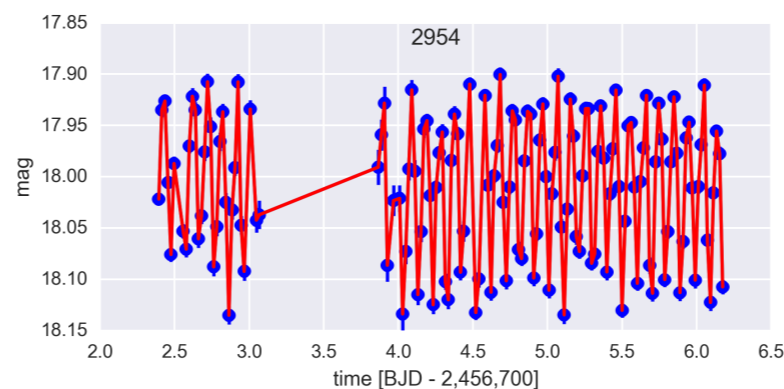
- **96** asteroids crossed the Nereid field
- **14** showed periodicities

• Goal: light curves \longrightarrow **shape reconstruction**

- $P_{\text{rot}} = \mathbf{2.78 - 14.63}$ h,
median: **4.74** h
- coverage: **0.05 - 14.63** d
median: **1.25** days

Neptune field

Szabó R. et al. in prep.
2016



Conclusions

- K2 provides a **unique way** to study Solar System objects
- **Long, uninterrupted, precise** photometry is unbeatable
- K2 can routinely observe in the **20-22.5 mag** range
- Covering the TNOs' path close to the stationary point only (where the movement is slowest) requires a **modest number of pixels**
- A few bright (20-22 mag) TNOs available per campaign
- **Main-Belt asteroids** and **moons** are also cool!



Thank you