

Debris & Formation

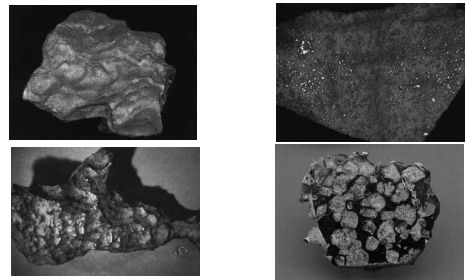
Meteors: Some Definitions

- **Meteor** - a bright streak of light observed when a piece of interplanetary debris enters a planetary atmosphere (also known as a 'shooting [or falling] star')
- **Meteoroid** - the aforementioned debris
- **Meteorite** - any surviving debris that reaches the planetary surface

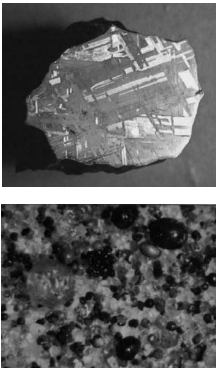
Interplanetary Debris

- Meteoroids/meteorites fall into two categories
 - **Stony** - silicate (rock) composition
 - **Iron** - iron/nickel composition
 - **Composite** - mixtures of silicate/metals
- Origin
 - Fragments of larger bodies
 - Size/age of parent bodies determined by crystallization patterns in rock/iron

Meteorite/Meteoroid Types



What's Inside



- Widmanstätten - metal 'grains' that are a result of heating
- Chondrites/Chondrules
 - Spherical bits of materials - never been heated

Process

- Bits of debris (with $v > v_{esc}$) passes too close to Earth's atmosphere
- Atmospheric resistance (aerobraking) slows debris to below v_{esc}
- Debris falls into atmosphere. Resistance elevates temperature -> fireball

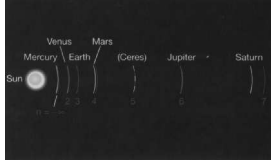
1766 – Titius-Bode Law

- Numeric sequence to determine planet location

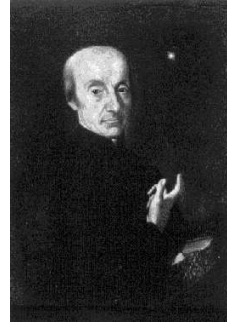
$$a_n = (0.4 + 0.3 \cdot 2^{n-2}) \text{ AU}$$

- Start with 0.4 AU (distance to Mercury)

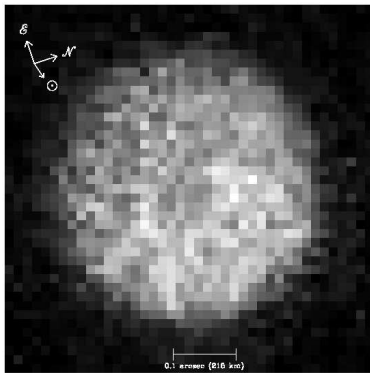
- 0.3 (to Venus) = 0.7 AU
- 0.3 (Terra) = 1.0 AU
- 0.6 (Mars) = 1.6 AU
- 1.2 (?) = 2.8 AU
- 2.4 (Jupiter) = 5.2 AU
- 4.8 (Saturn) = 10.0 AU
- 9.6 (Uranus) = 19.6 AU



1801: Discovery of Ceres



- Jan 1 - Giuseppe Piazzi (1746-1826), in an attempt to verify Titius-Bode, finds the 'fifth planet' between Mars and Jupiter



Ceres

- $D = 910 \text{ km}$
- $a = 2.77 \text{ AU}$
- $P = 4.61 \text{ years}$

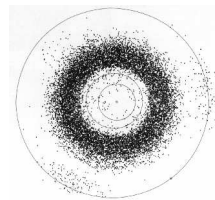
Asteroids/Minor Planets

- Silicate/metallic, ~1-1000km diameter
- Stable orbits around Sol
- Most located in belt between 2.1-3.3 AU
- Many located further out/closer in
 - Near Earth asteroids
 - highly elliptical orbits
 - Trojan asteroids
 - Lagrange/libration points



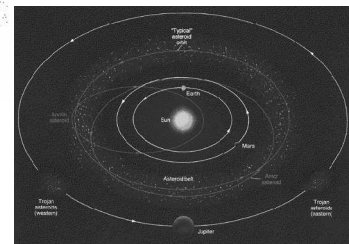
Asteroid Composition

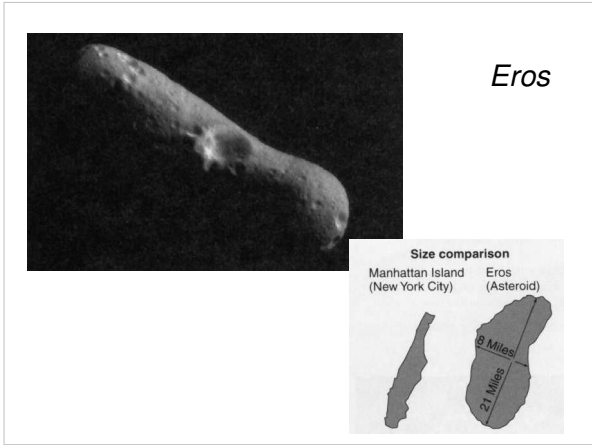
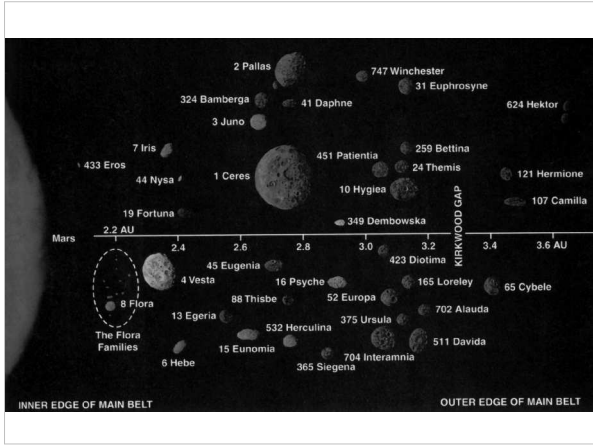
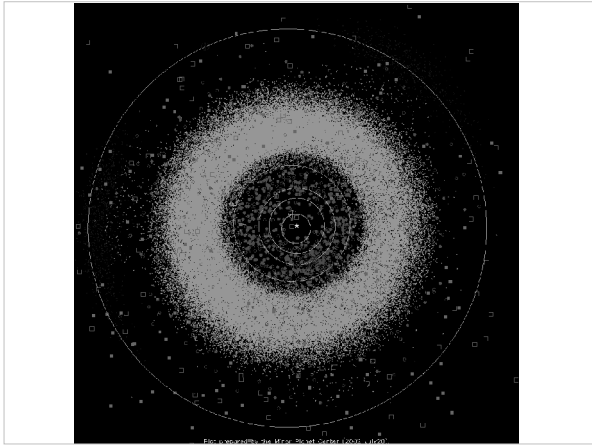
- **C-type** - dark, low albedo
- **S-type** - silicates, olivine
- **M-type** - metallic iron/nickel
- **V-type** - silicates, basalts



Asteroid Orbits


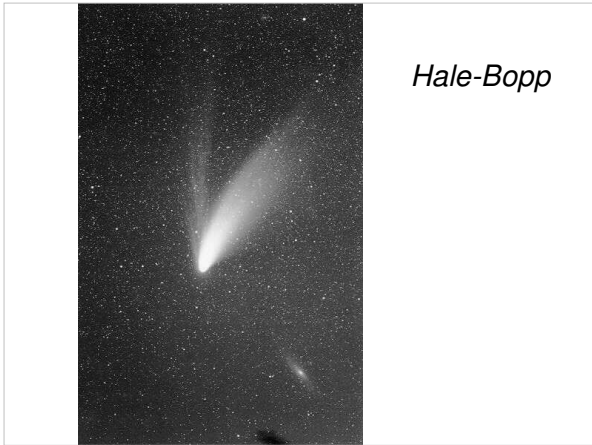
- Apollo - Earth crossing
 - Perihelion > 1 AU
- Aten - Earth crossing
 - Perihelion < 1 AU
- Amor - Mars crossing
- Trojan - Jupiter lagrange



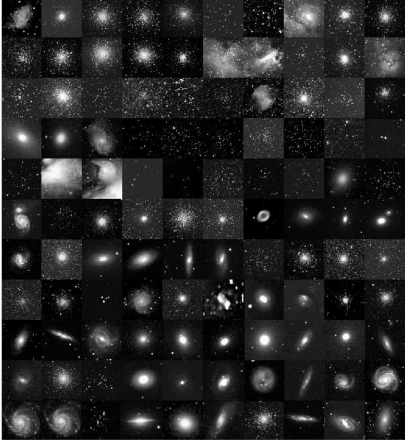


Comets

- Edmund Halley (1705) predicted that the orbits of comets observed in 1531, 1607, 1682 were the same -and 76 years apart. He predicted a return of the comet in 1758.
- Comet was observed Christmas night of that year, 16 years after Halley's death.

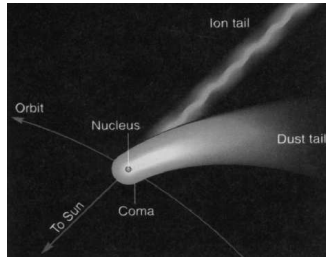
Not Comets.



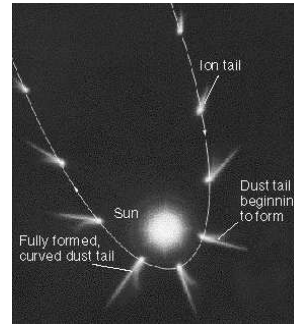
- Charles Messier (1730-1817) -comet hunter.
- Made a list of 101 things that were *not* comets.

Comet Structure

- Nucleus < 10 km
 - H₂O, CO, CO₂, Formaldehyde
- Coma ~10⁶ km
- Tails ~ 10⁸ km
 - ion (charged)
 - dust (uncharged)



Tail Formation



- **Dust tail**
 - uncharged
 - white/yellow
- **Plasma tail**
 - ionized CO
 - blue

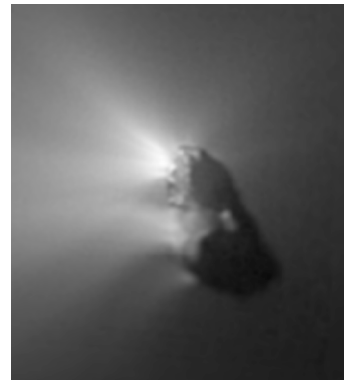
Comet Hyakutake



Comet Hyakutake - C/1996 B2
Hubble Space Telescope - Wide Field and Planetary Camera 2



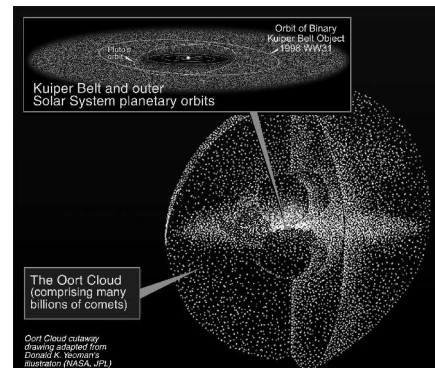
Halley's Nucleus



- Jets on sunlit side
- Surface of fragmented material
- Interior ~200K
- Surface ~350K

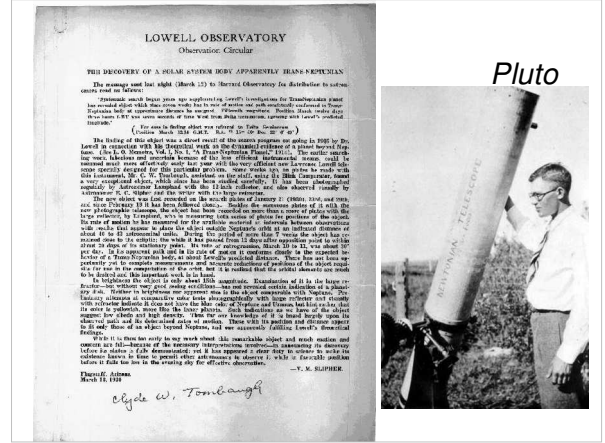
Cometary differences & Sources

- | | |
|---|---|
| <ul style="list-style-type: none"> • SHORT PERIOD – <ul style="list-style-type: none"> - $i \leq 30^\circ$ - $P < 500$ yrs | <ul style="list-style-type: none"> • LONG PERIOD – <ul style="list-style-type: none"> - $0 \leq i \leq 180^\circ$ - $P > 500$ yrs |
| <ul style="list-style-type: none"> • Edgeworth-Kuiper Belt - flattened disk of icy bodies at 30 to 50 AU <ul style="list-style-type: none"> - Comets would be dislodged by gas giants' gravity - comets with periods of 70-1000 years | <ul style="list-style-type: none"> • Oort cloud - spherical swarm of icy bodies at 10,000 to 75,000 AU <ul style="list-style-type: none"> - Comets formed when gravitational influence knocks snowball out of cloud |

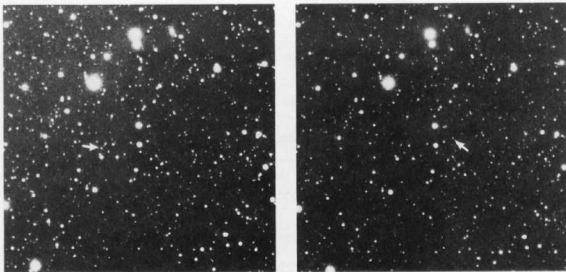


Oort Cloud gateway
Drawing adapted from
Donald K. Yeoman's
illustration (NASA, JPL)

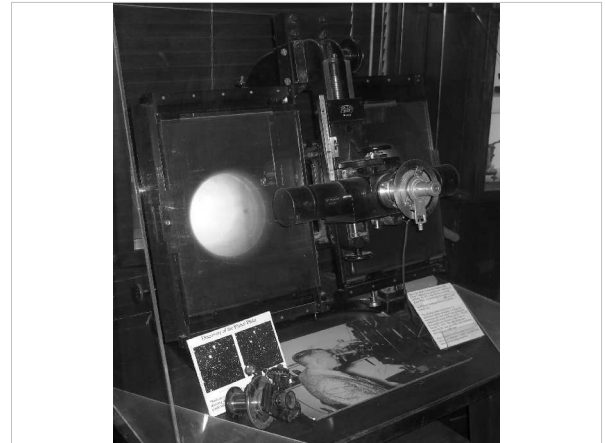
Planet P vs. Planet X



Needle in a Haystack



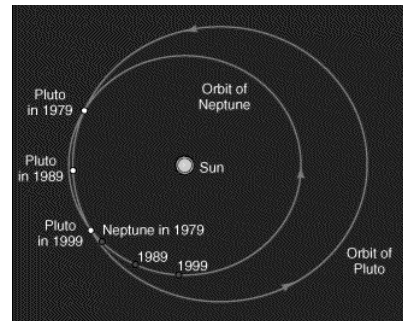
- L exposure - 01/23/30, R exposure 01/29/30
- Field ~ 6'38" x 6'38", displacement ~ 84"

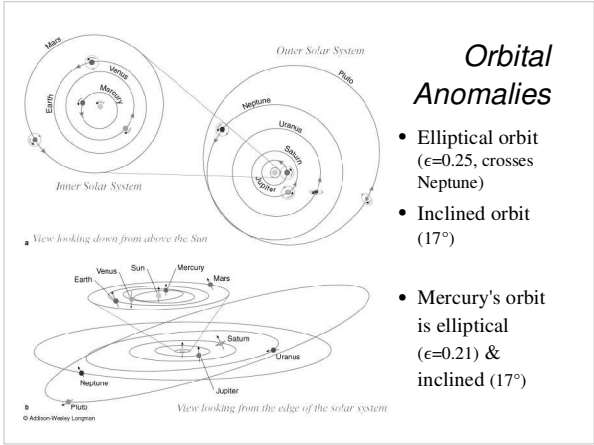


1930/03/13: Pluto Discovered

- Semimajor axis: 39.5 AU
- Sidereal year: 248.5 years
- $0.0022M_{\oplus}$, $0.18R_{\oplus}$, $\rho \sim 2000 \text{ kg/m}^3$
- Rotation period: 6.4 mean solar days
- Axial tilt: 122°
- Orbital eccentricity: 0.248
- Inclination of orbit to ecliptic: 17°

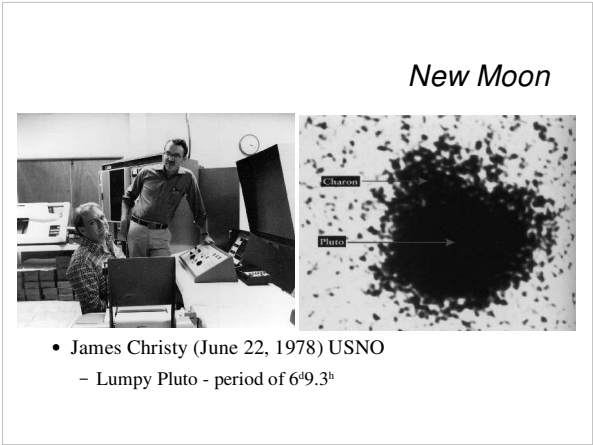
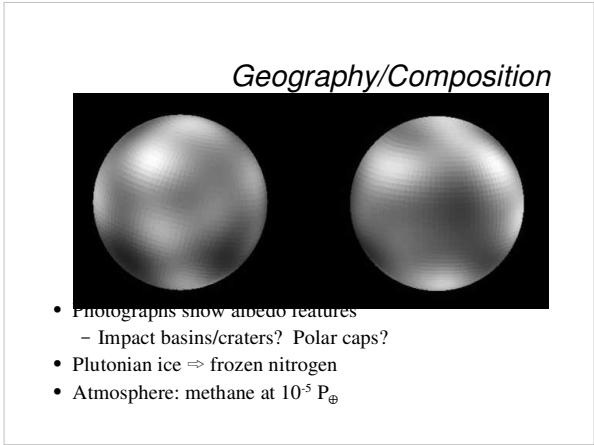
Pluto's Eccentric Orbit





Pluto's Shortcomings

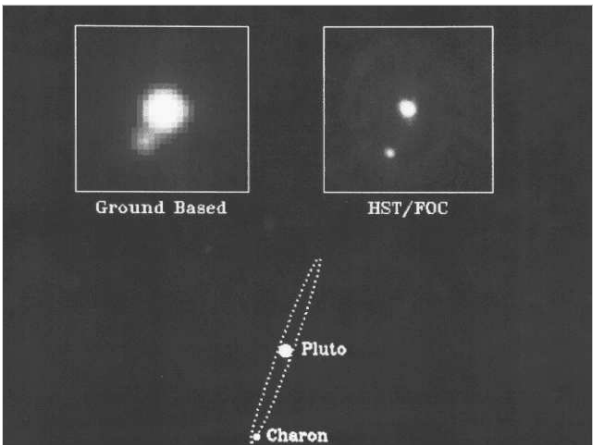
- While Pluto seems to be where Lowell predicted (more or less), there are problems.
 - Mass and size are too small (by a factor of 5 at least)
 - Highly elliptical orbit (crosses Neptune's)
 - Highly inclined orbit
 - Out of place for a terrestrial planet
 - Obviously not a jovian planet



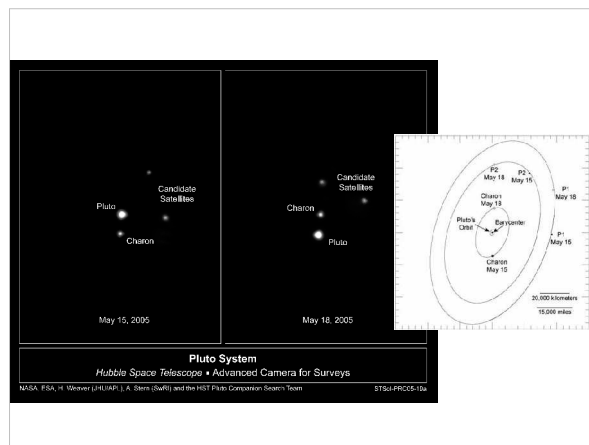
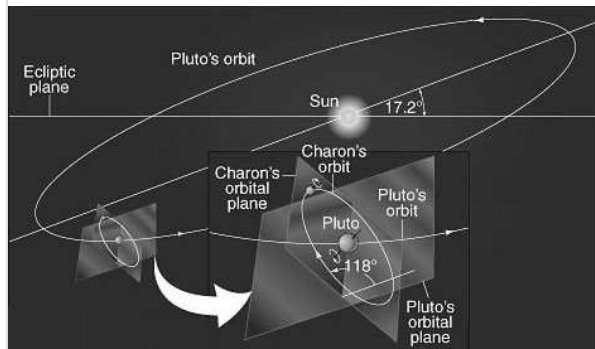
Pluto/Charon Binary System

- Distance from Pluto $\sim 1/20$ distance between Terra & Luna
- Diameter - 1270km (Pluto is 2320km)
- Mass - $1/10th M_E$
- Charon allowed determination of Pluto's mass
- Synchronous rotation (6.4 days)

Pluto, Charon, & USA Comparison



Orbit Detail

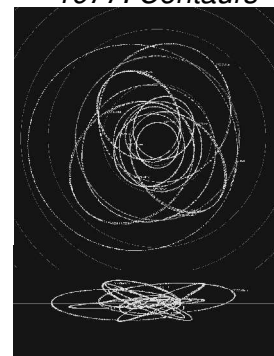


Pluto's Shortcomings

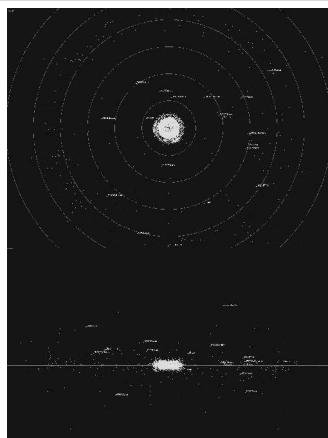
- Pluto's moon is too big for her
- Obviously not a jovian planet
- Neither very rocky or metallic for a terrestrial planet
- Orbit is highly inclined to the ecliptic
- **No answer for these...**

1977: Centaurs

- Charles Kowal (Palomar)
 - Chiron
 - '10th planet', asteroid
 - Developed 'tail'
- Centaurs:
 - Carbonaceous chondrites
 - Ices
 - Diameters ~30-300km
 - Highly elliptical, inclined orbits

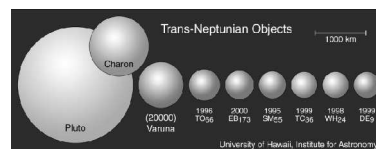
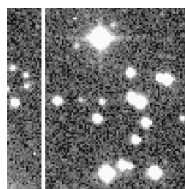


1992 QB₁

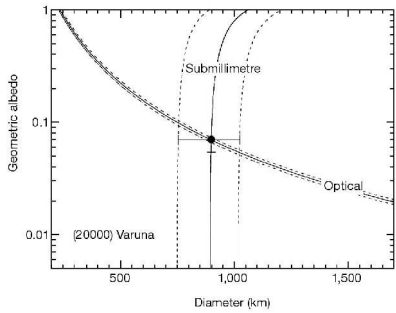


- Cubewanos
- Plutinos
- KBOs
- TNOs
- SDOs

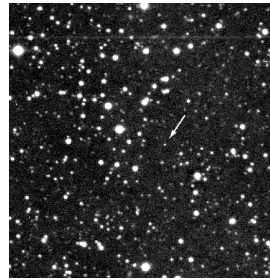
(2000 WR₁₀₆) 20000 Varuna



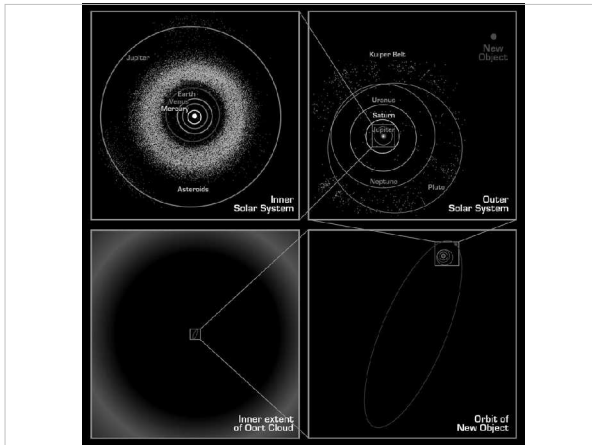
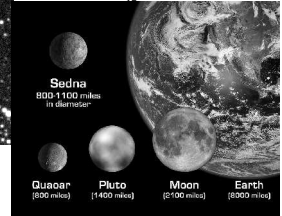
Determination of Size



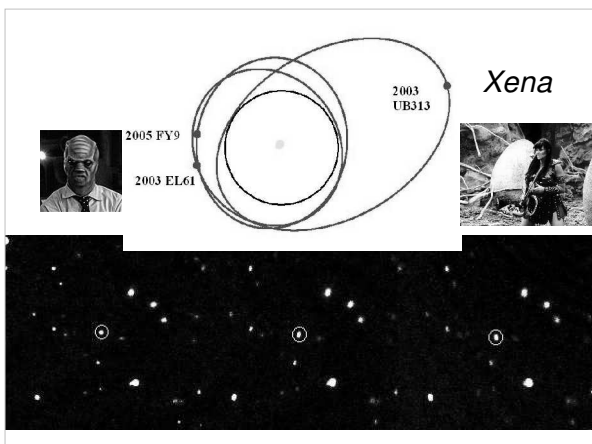
Challengers

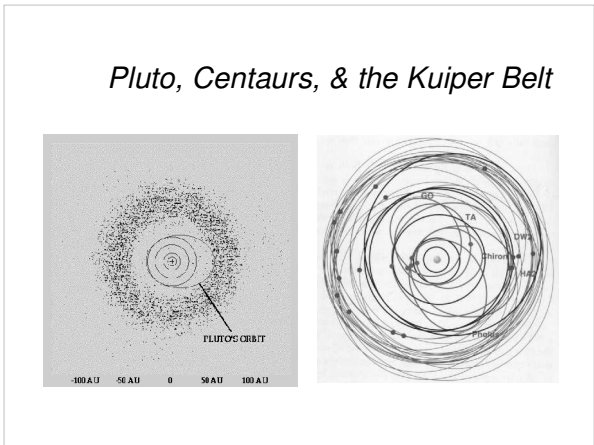
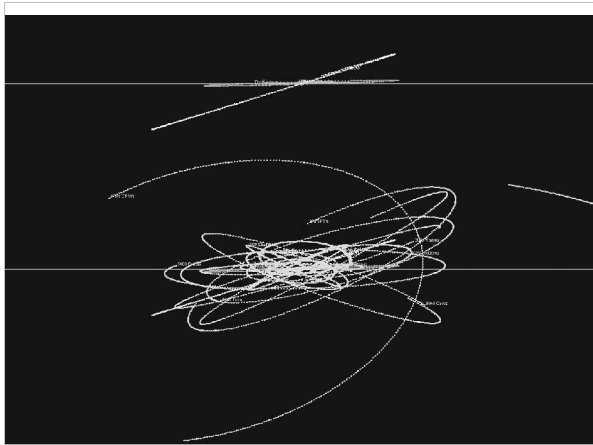
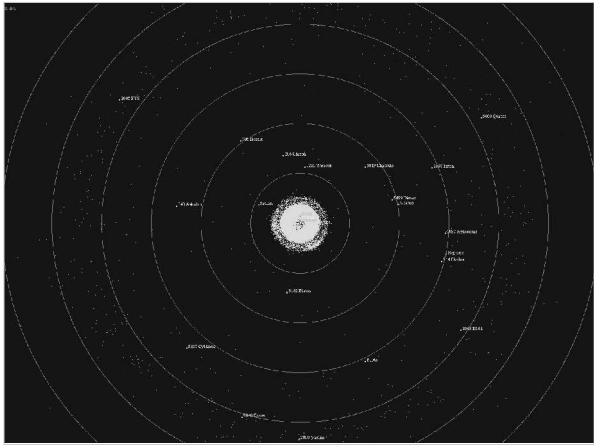
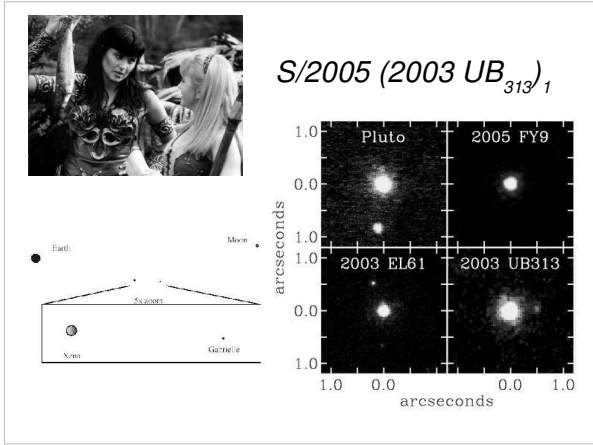
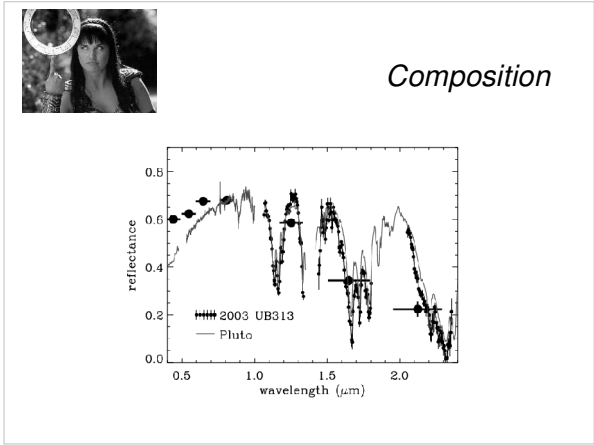


- (2001 KX₇₆) 28978 Ixion
- (2002 LM₆₀) 50000 Quaoar
- (2003 VB₁₂) 90377 Sedna



2003 UB₃₁₃ (Xena)





Remember our Planetary System Model Criteria?

- Each planet is relatively isolated in space
- Orbits of planets are nearly circular
- Orbits of planets lie in the same plane
- Direction of orbits is same as Sol's rotation
- Planetary rotation is in the same direction as Sol's rotation

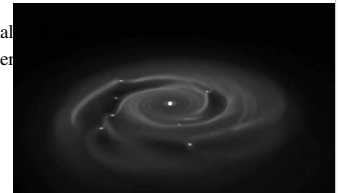
- Most moons revolve in the same direction as planetary axial rotation
- Planetary system is highly differentiated
- Asteroids are very old and share properties unlike terrestrial or jovian planets
- Comets are icy, primitive fragments that do not orbit in ecliptic and reside far from the Sun

The Solar Nebula

- Est. 4.6 billion years ago
- 10-100 million years to form planets
 - isotope abundances
- mass $< 0.1 M_{\oplus}$
 - most mass blown away by stellar wind
 - $\rho_{avg} \sim 10^{-7} \rho_{air, \oplus}$

Condensation

- Heat produced by Kelvin-Helmholtz contraction
 - Temps from several thousand K near core to 100K at 10 AU (Saturn)
- Particles condense from cooling gas
- **Differentiation**
 - stable materials (metal
 - hydrogen, helium later



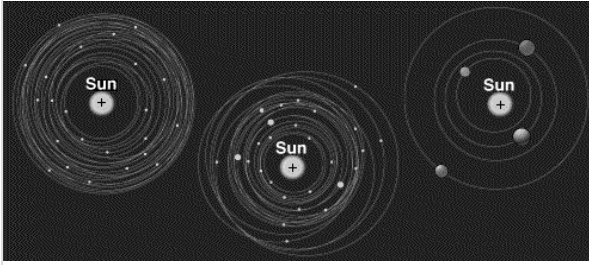
Planetesimal Formation

- Condensed matter (grains) slowly stuck together
- **Accretion** - as pieces stuck together, gravitational force increased and attracted more pieces
- **Planetesimals** - larger accumulations of matter, from 10^3 to 10^5 m (asteroids)
- Planetesimals likely aided condensation process

Terrestrial Planet Formation

- Formation began near 1600K
 - metals, silicates, very little water/gases
- Surfaces of Mercury, Mars highly suggestive of planetesimal collisions
- **Underrepresented** in light material
 - Temperatures so warm that gases condense
 - Surface gravities too small to retain
 - Gases likely 'blown away' by protostellar winds

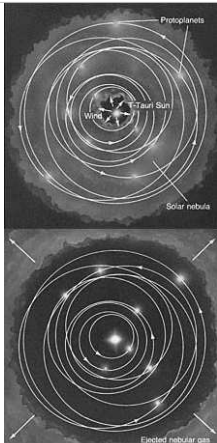
Planetesimal Aggregation



Jovian Planet Formation

- Formation began almost immediately
 - Cooler temperatures in outreaches allowed immediate condensation of gases
 - Lighter gases condensed around cooler bits of iron/silicates
 - As cooler planetesimals swept through nebular gas, warmer gas condensed on the planetesimals

Cleanup



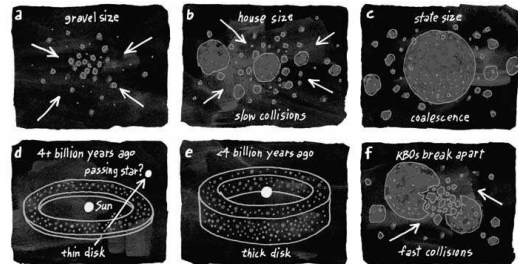
- Light, ice-based planetesimals probably ejected outward to form Oort Cloud, Kuiper Belt
- Non-aggregate nebular gas blown away from solar system

Cleanup

- As planets form, the original elliptical orbits of the planetesimals 'average' out to nearly circular orbits.
- Eventually, most of the material near the newly-formed planets is 'swept' up.
 - several of the terrestrial planets still bear the scars of the last impacts?

Residue

- In areas, some of the proto-matter remains - never accreted into a larger body.
 - Asteroids
 - Kuiper Belt/Scattered Disk
 - Oort Cloud
- These objects generally have more elliptical/inclined orbits



Review

- Condensation theory accounts for all of our observations.
 - Planets' orbits are circular, in same plane, and in same direction as Sol's rotation due to nebula's shape and rotation.
 - Rotation of planets and moon systems due to tendency of small eddies to inherit overall rotation of nebular disk

- Growth of planetesimals results in widely-spaced orbits
- Heating of nebula, solar ignition result in differentiation
- Debris from accretion-fragmentation account for asteroids/comets

Catastrophism: Exceptions Explained

- Mercury's large nickel-iron core
- Venus' low rotation rate
- Luna's orbit, size
- Mars' asymmetry/lack of atmosphere
- Uranus' axial tilt
- Uranus' moon Miranda
- Triton's retrograde motion
- Pluto-Charon system

