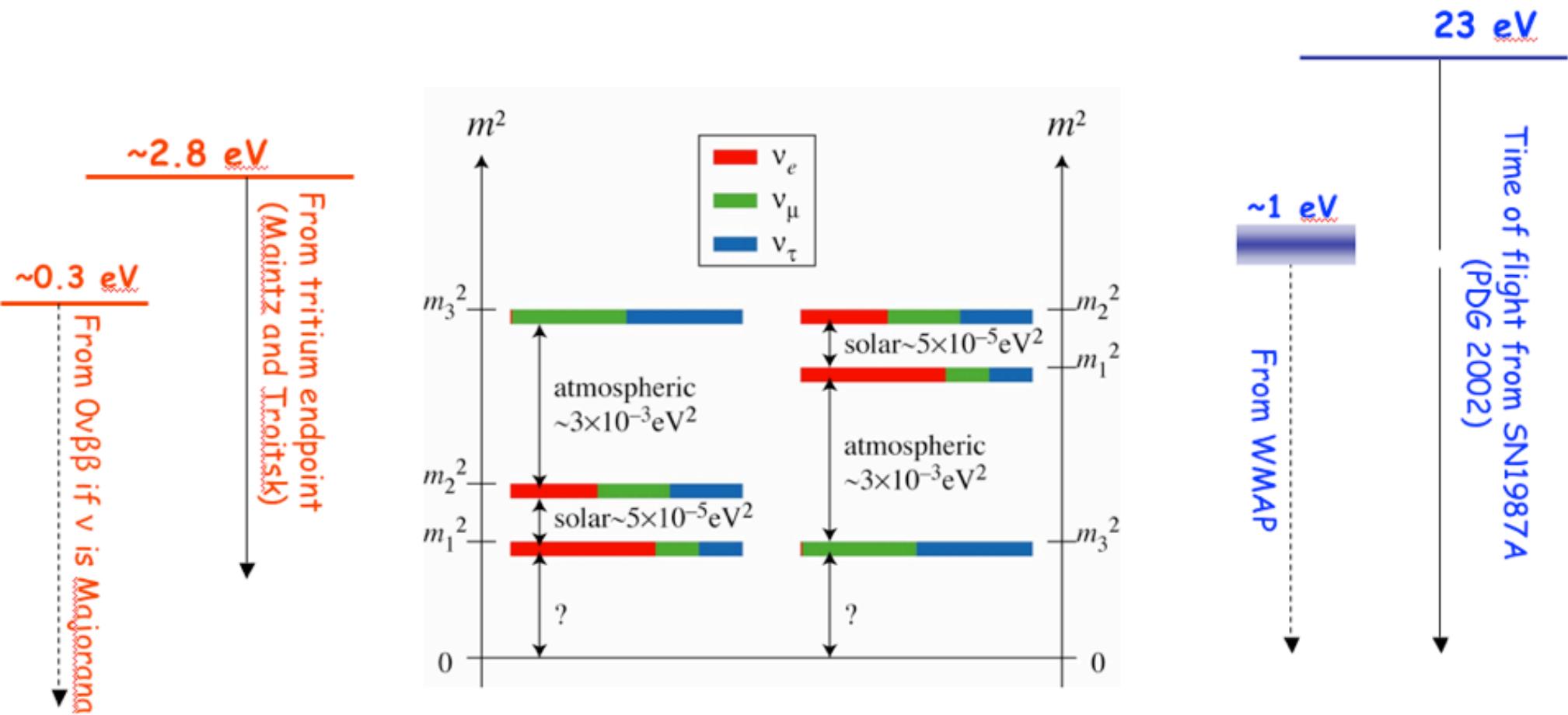


# Status of $\beta\beta$ -decay in Xenon

(or *Back to the Future*)

Roland Lüscher  
CCLRC/RAL

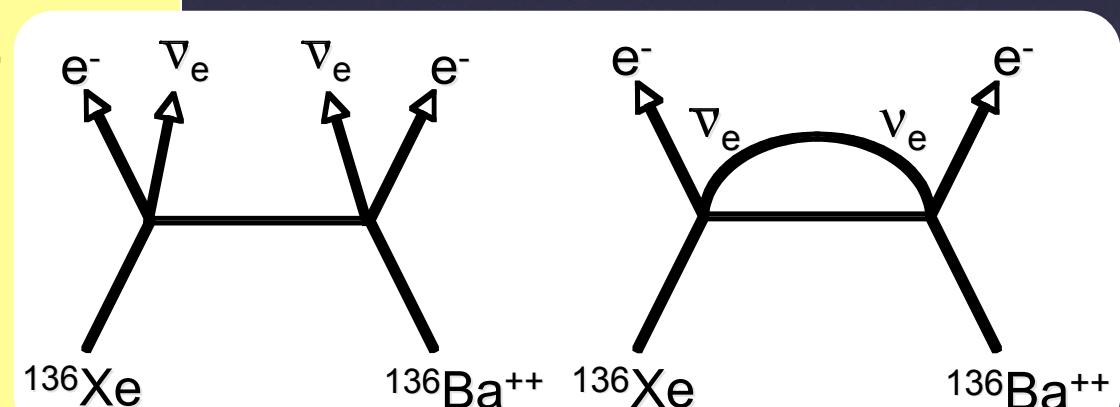
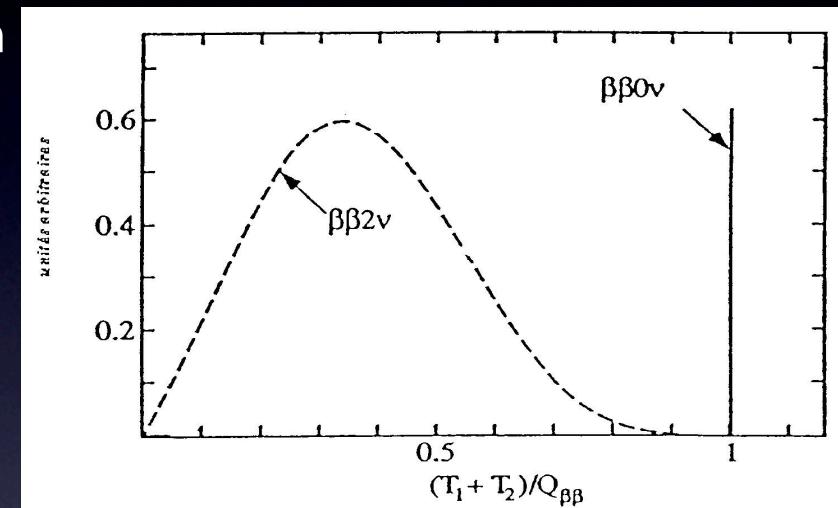
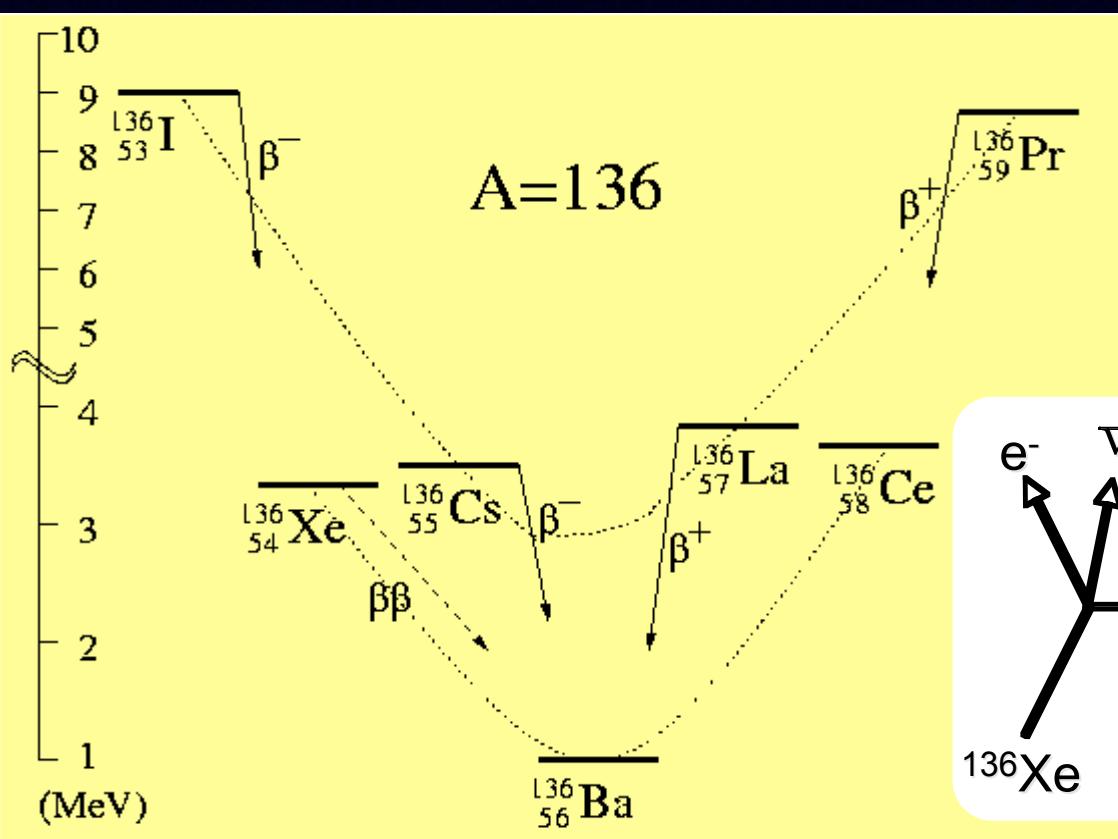
# The most important question in neutrino physics: Neutrino mass scale



ie. need a  $\sim 10\text{-}200 \text{ meV}$  sensitivity  
 $\sim 1\text{e}27\text{-}1\text{e}28 \text{ yr}$  in  $^{136}\text{Xe}$

# $\beta\beta$ -decay

- Second-order process, only observable when  $\beta$ -decay is energetically forbidden (or strongly suppressed by angular momentum differences)



# $\beta\beta$ -decay in Xenon

- 3 candidates to look at:
  - $^{136}\text{Xe}$ , with  $E_0 = 2.48 \text{ MeV}$ , 8.9% natural abundance
  - $^{134}\text{Xe}$ , with  $E_0 = 0.85 \text{ MeV}$ , 10.4% natural abundance
  - $^{124}\text{Xe}$ , with  $E_0 = 0.82 \text{ MeV}$ , 0.1% natural abundance  
( $\beta\beta++$  decay !!)
- neutrinoless- $\beta\beta$  rate, is going with  $E_0^5$

$$\langle m_\nu \rangle^2 = \left( T_{1/2}^{0\nu\beta\beta} G^{0\nu\beta\beta}(E_0, Z) \left| M_{GT}^{0\nu\beta\beta} - \frac{g_V^2}{g_A^2} M_F^{0\nu\beta\beta} \right|^2 \right)^{-1}$$

- World yearly production of enriched  $^{124}\text{Xe}$  is ... 5 liters!
- $\rightarrow$   $^{134}\text{Xe}$  and mainly  $^{136}\text{Xe}$

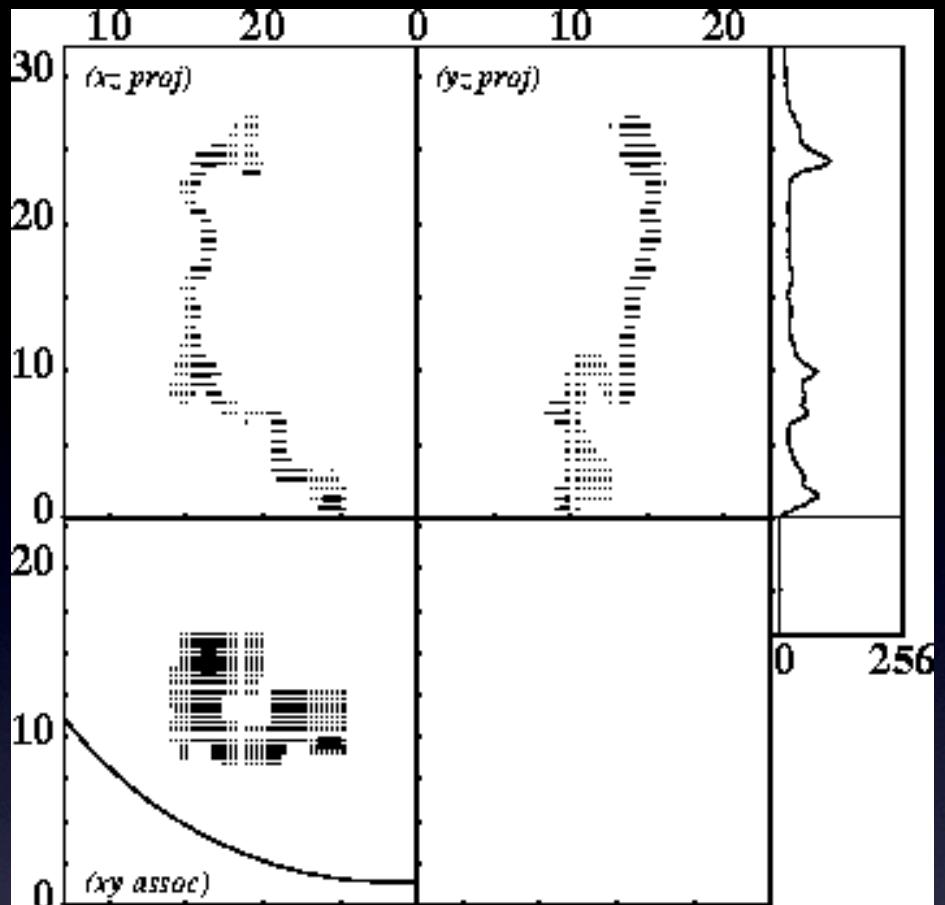
# Why Xenon?

- Low background - once  $^{85}\text{Kr}$  removed  
(and can be cleaned without ‘re-crystallising’)
- ‘easy’ enrichment  
(Noble gas = no chemistry needed, centrifuging rate  $\sim \text{g/s}$ ).
- Scintillation properties enables **Pulse-Shape Discrimination**  
(PSD) between background (electron recoil) and nuclear recoil events.
- **Ionisation** properties ...
- Scale up possibilities - up to 1 or 10 tons  
(World yearly production: 36t)

# Typical Problems with Xe

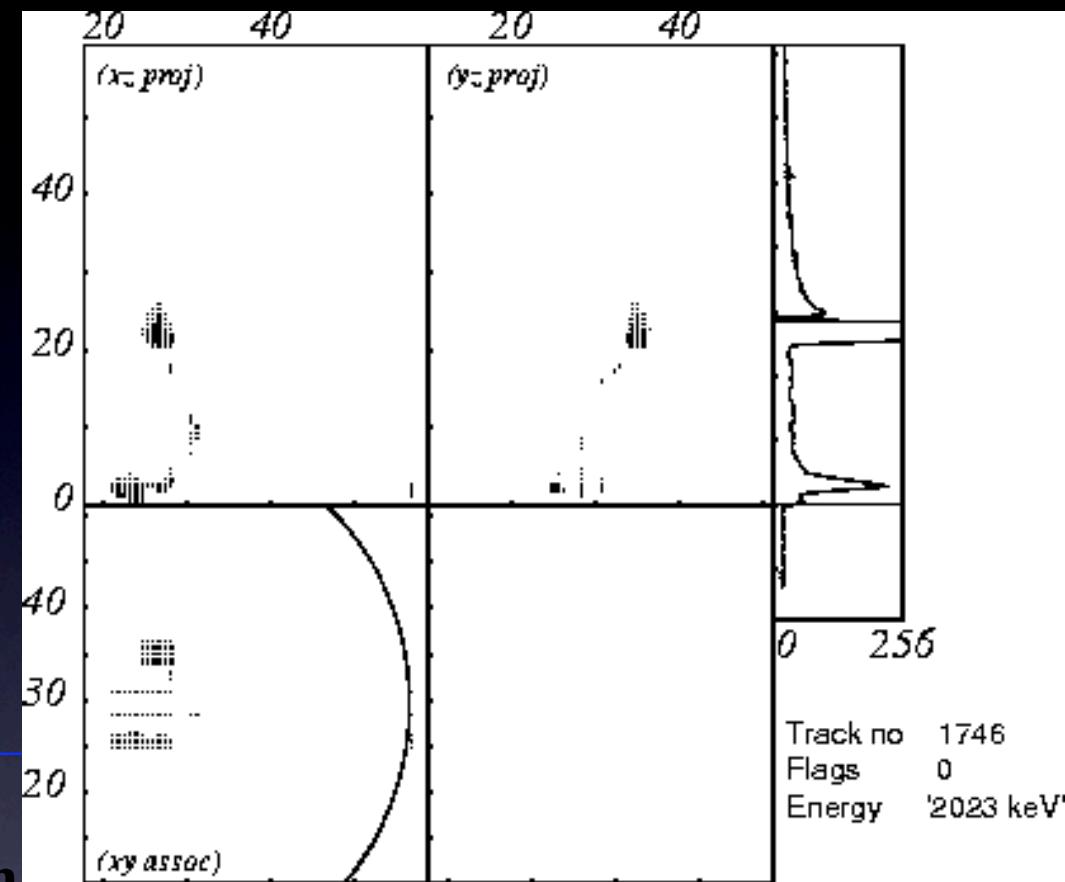
- Energy Resolution - not a solid state detector
- Gas-Xe allows tracking, but large mass takes large volume
- Liquid-Xe is very dense but loses tracking capability

# Gas Xenon

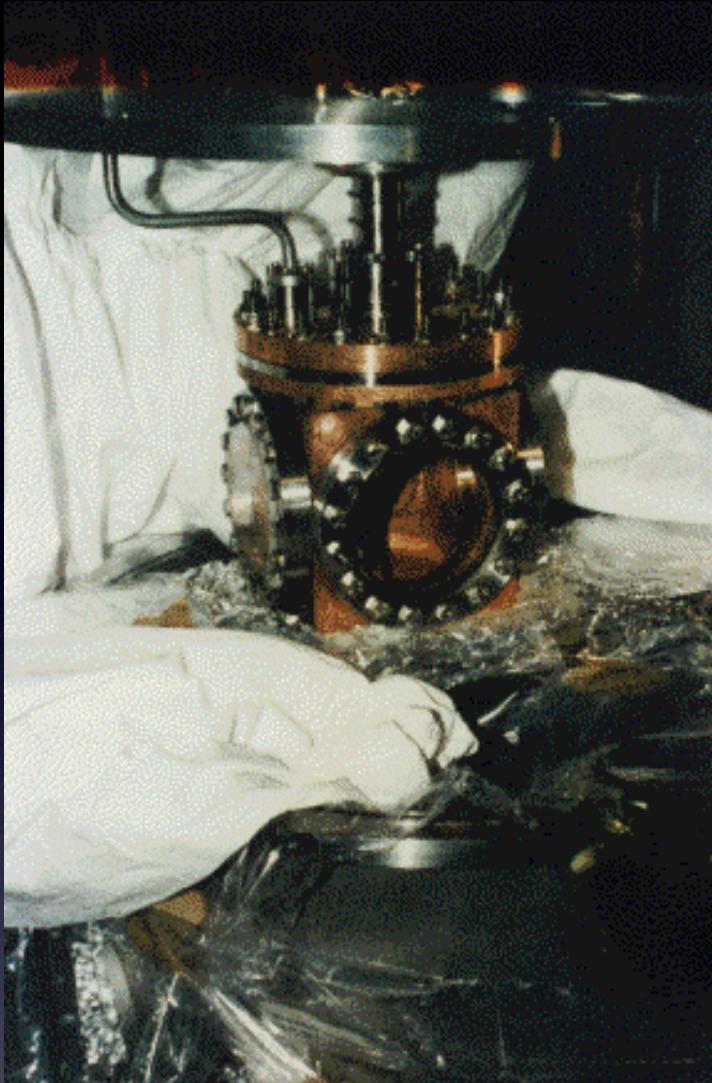


**Advantage:**  
Tracking background discrimination

**Disadvantage:**  
Large detector for small source mass  
Limited energy resolution  
( $\sigma(E)/E = 2.8\%$  at 2480keV)



**The Gotthard TPC**  
Volume: 180l  
3.3 kg of  $^{136}\text{Xe}$  (62.5% enrichment)



## The DAMA LXe chamber

Volume: 1.5l

4.5 kg of  $^{136}\text{Xe}$  (68.8% enrichment)

# Liquid Xenon

Gotthard:  
DAMA:

### Advantage:

Higher density -> smaller detector  
Potentially Improved energy resolution  
(Zeplin1:  $\sigma(E)/E = 2.3\%$  at 2480keV  
Dama:  $\sigma(E)/E = 8.8\%$ )

### Disadvantage:

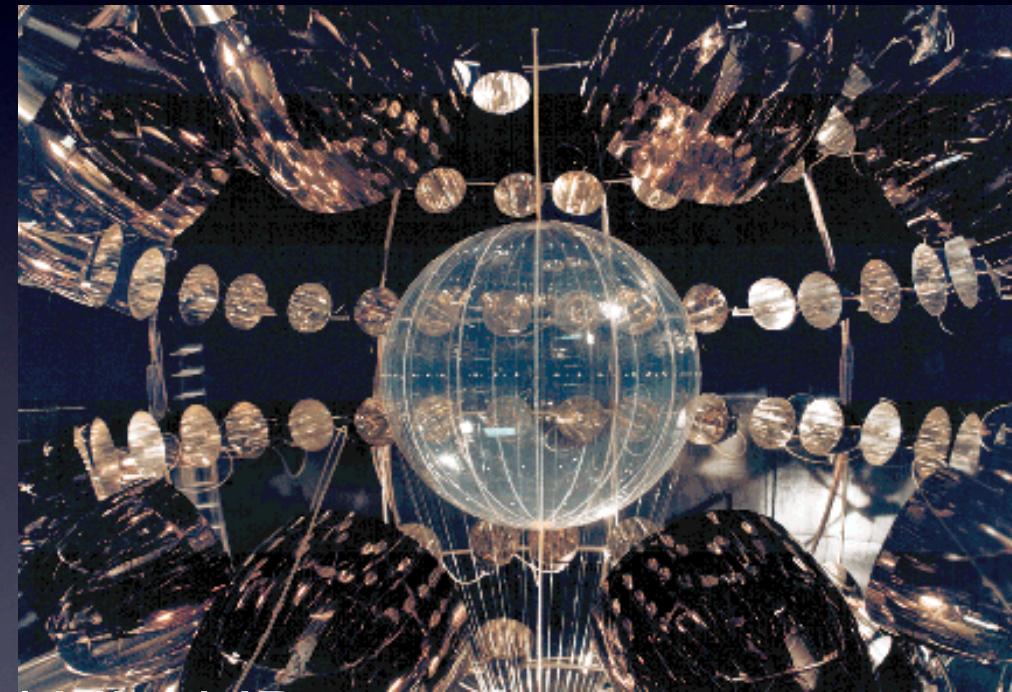
‘Only’ pulse shape discrimination  
Detector response non-uniformity?

# Current status

- Gotthard GXe TPC  
 $T_{1/2} > 4.4 \times 10^{23}$  yr,  $\langle m \rangle < 1.8\text{-}5.2$  eV  
Luscher, PLB, 434 (1998) 407
- DAMA LXe scintillation detector  
 $T_{1/2} > 1.2 \times 10^{24}$  yr,  $\langle m \rangle < 1.1\text{-}3.1$  eV  
Bernabei, PLB, 546 (2002) 23
- ...

# Dissolved Xenon

- The third solution: dissolve Xenon in a large scintillation detector (Raghavan, PRL 72 1411 (1994))
- BOREXINO: 300t scintillator  
1.565t of Xe (enriched to 80%)  
sensitivity:  $1.14 \times 10^{27}$  yr  
backgnd: 0.003 /kev/kg/yr  
(Caccianiga, Giammarchi,  
Astropart Phys 14 15 (2000))

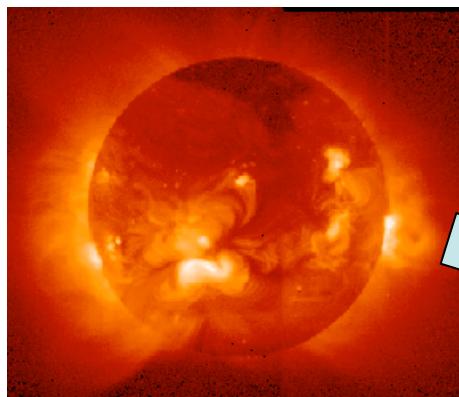


- Also evoked for SNO and KAMLAND ...  
... but unlikely to happen in either 3!

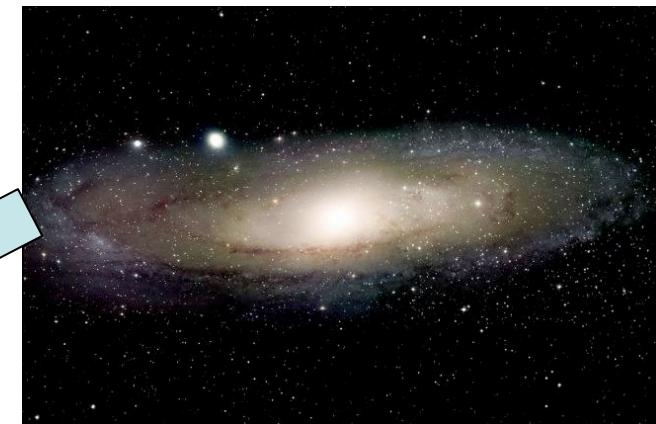
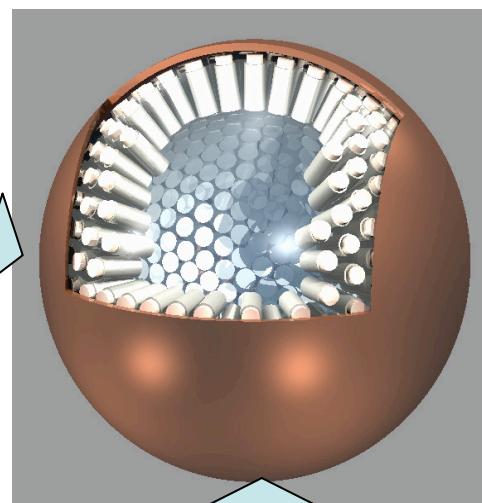
# XMASS

Multi purpose low-background experiment with liq. Xe

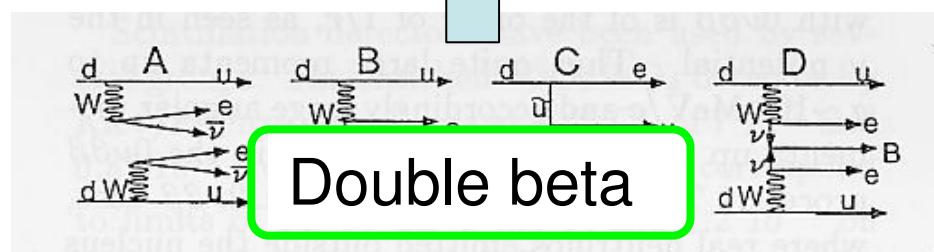
- Xenon **MASS**ive detector for solar neutrino ( $\text{pp}/^7\text{Be}$ )
- Xenon neutrino **MASS** detector ( $\beta\beta$  decay)
- Xenon detector for Weakly Interacting **MASS**ive Particles (DM search)



Solar neutrino

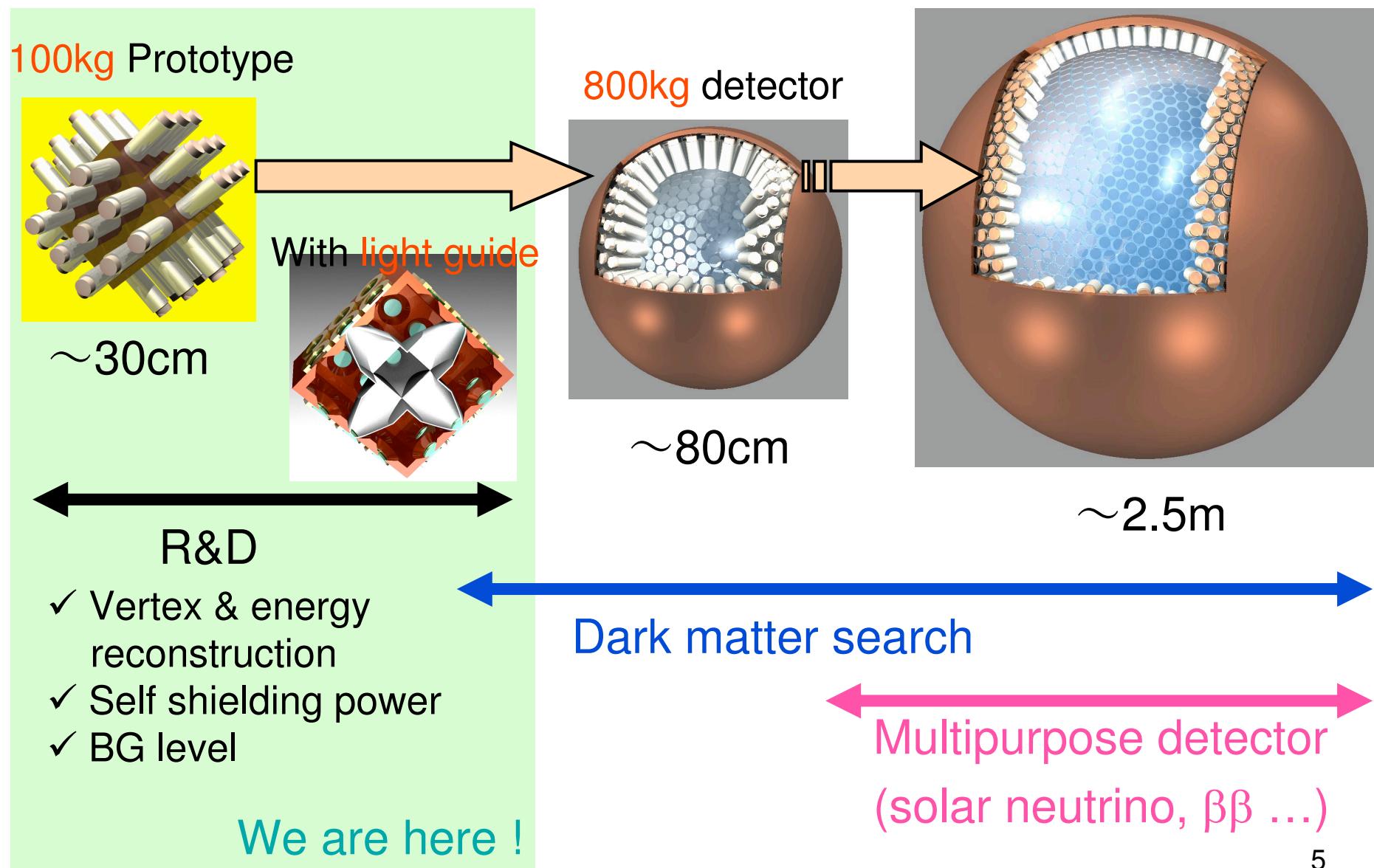


Dark matter

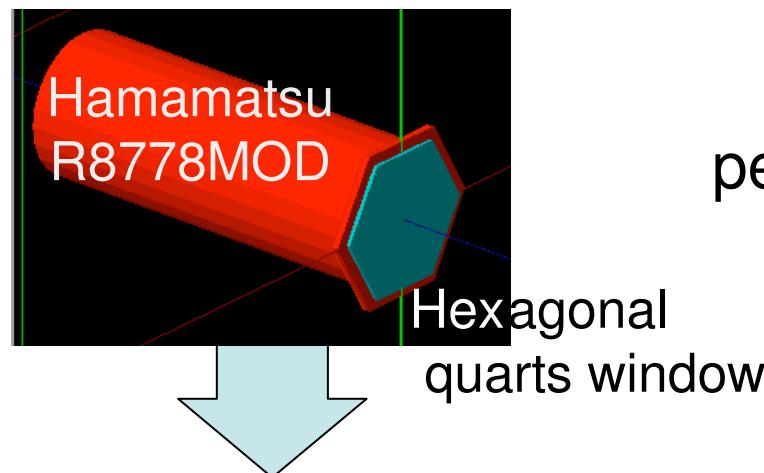


2

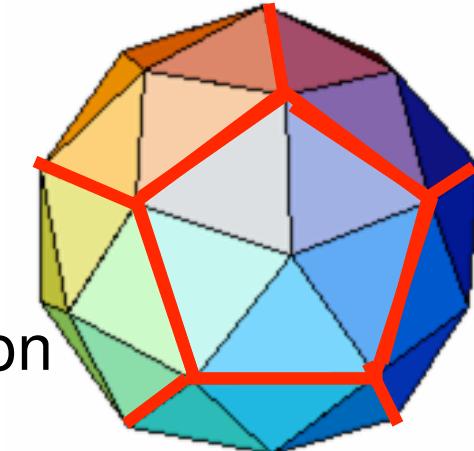
# ➤ Strategy of the scale-up



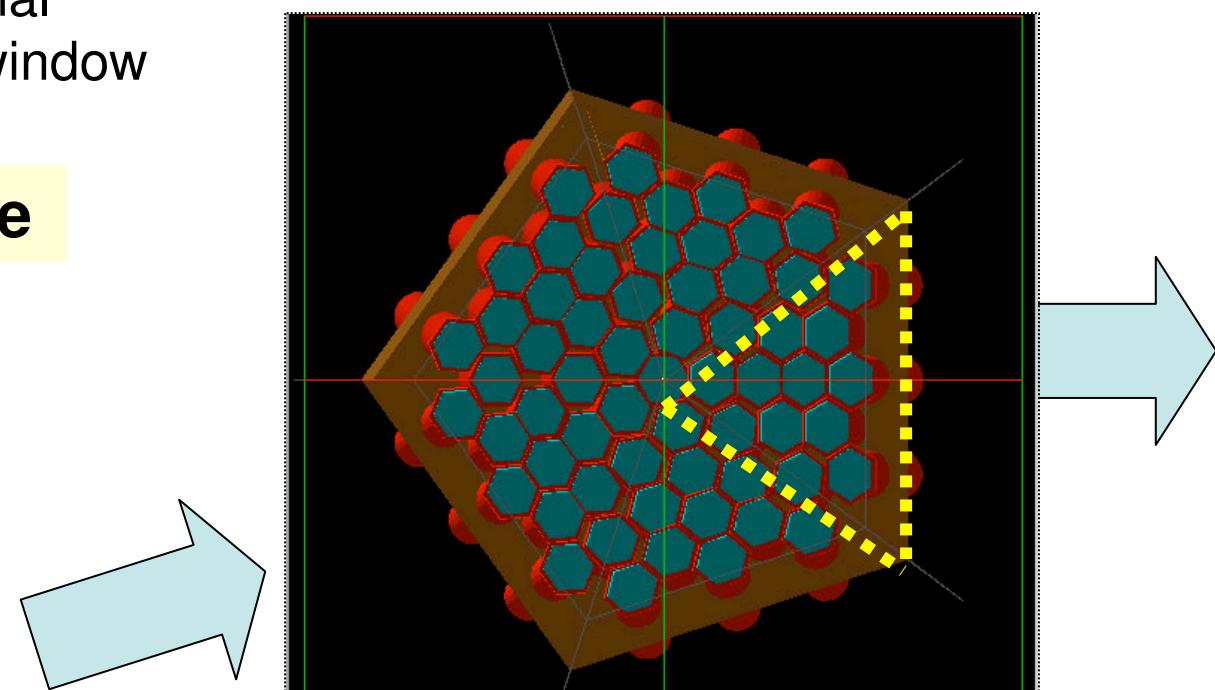
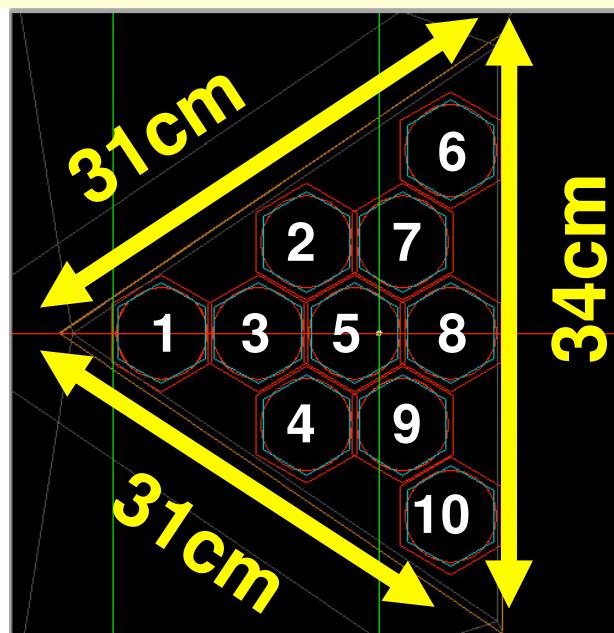
## ➤ Structure of 800 kg detector



12 pentagons /  
pentakis dodecahedron

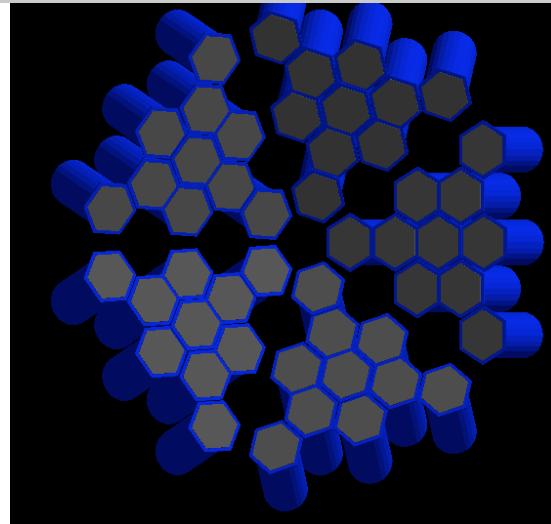
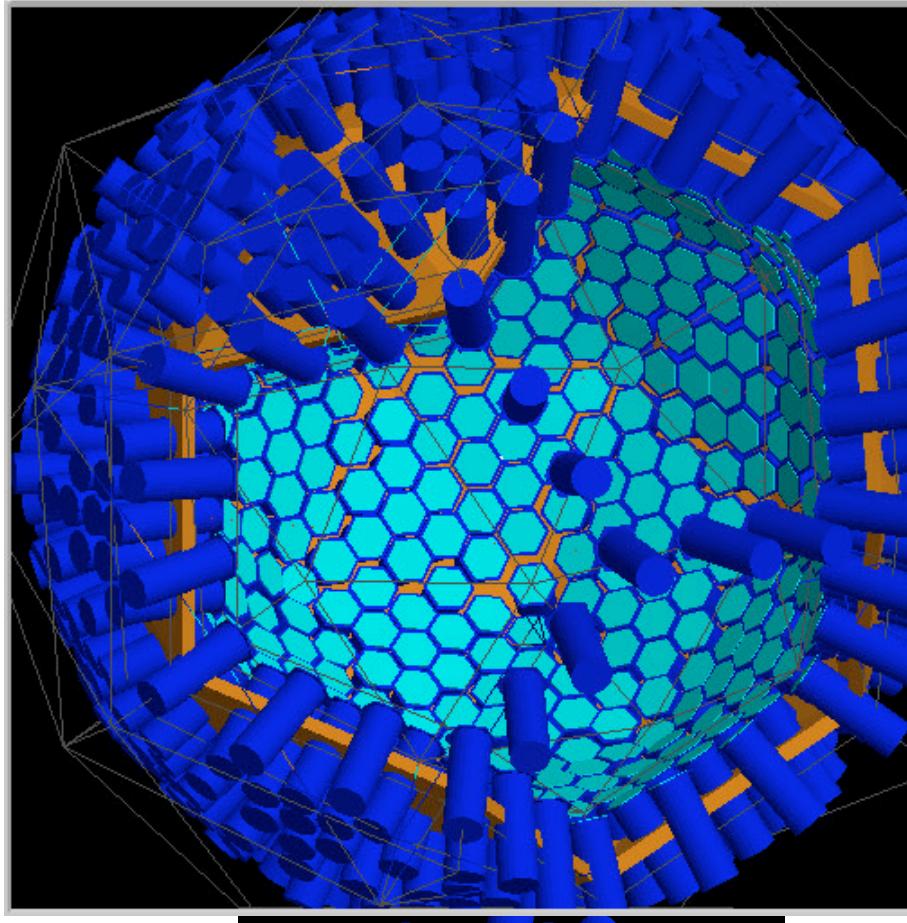


10 PMTs per 1 triangle

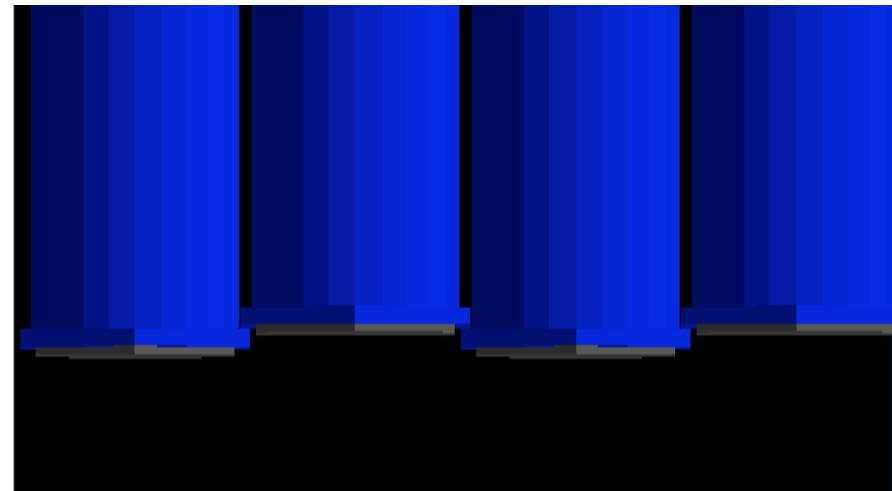


5 triangles make pentagon

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- Total **812 hex PMTs**  
(10PMTs/triangle  $\times$  60 +  
212 @gap)  
immersed into liq. Xe
- **~70%** photo-coverage
- Radius to inner face **~44cm**

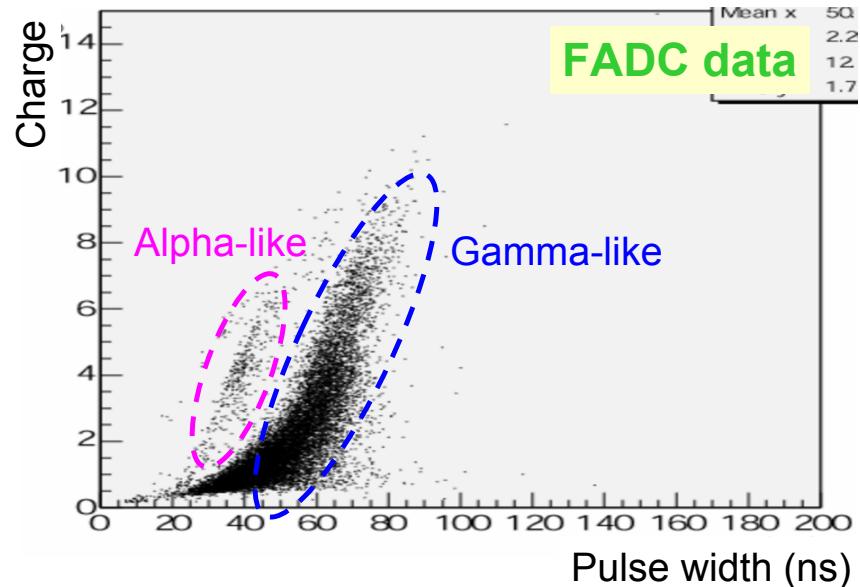


Each rim of a PMT overlaps  
to maximize coverage

12

## Alpha vs Gamma separation

A  
P



Alpha-gamma separation by using FADC wave form would be possible (under further investigation)

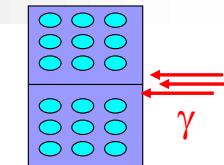
August 18, 2004

Y.Takeuchi @ICHEP04 in Beijing

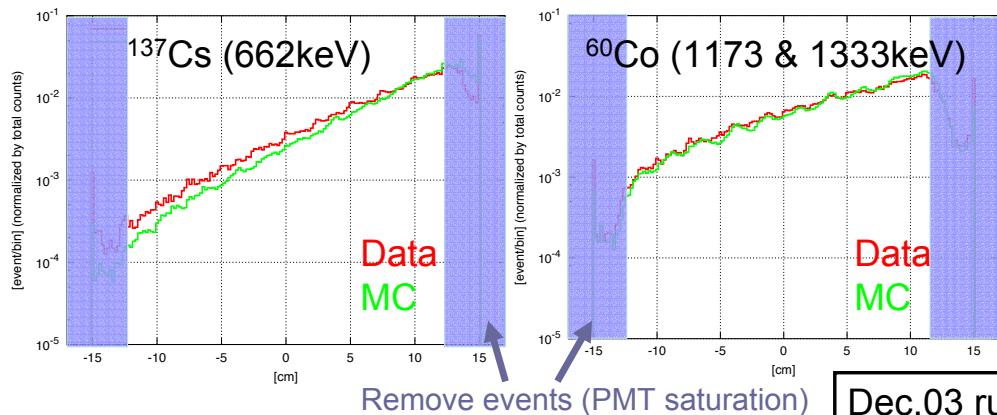
# 100kg Prototype measurements

- Good position reconstruction
- background 11 c/kev/kg/yr !!!!

## Self shielding performance



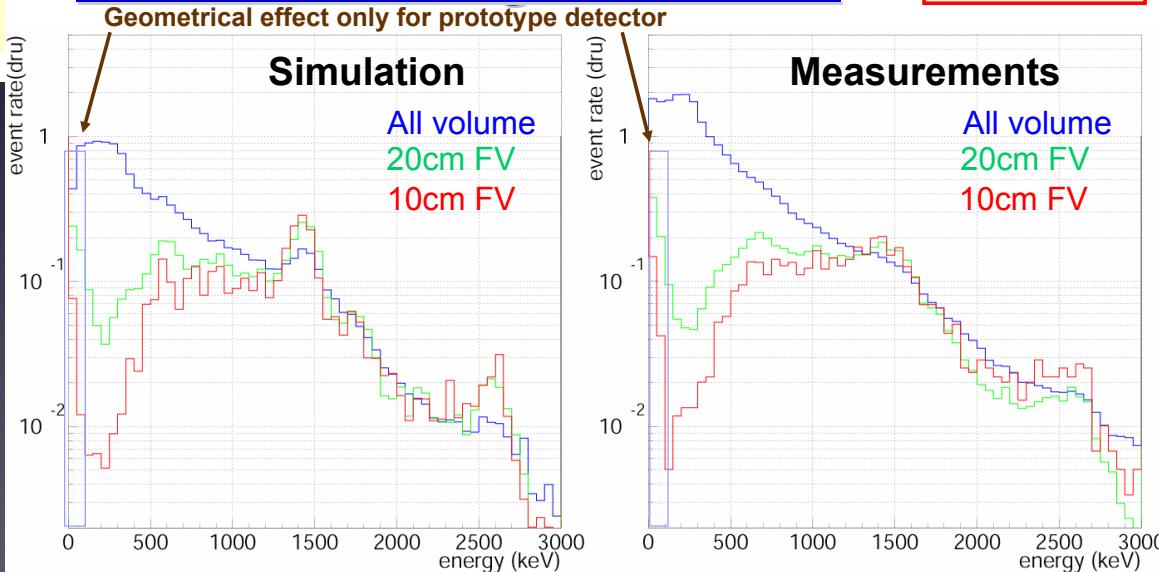
Reconstructed vertex position of collimated source runs



Aug.04 run  
Preliminary

## Measured background level

Geometrical effect only for prototype detector



- Self shielding works
- Good agreement with expectation (< factor 2)

# XMASS prospects

100kg prototype: 11 c/kev/kg/yr background

By comparison :

Zeplin1 had 0.05 c/kev/kg/yr at 2.5MeV

Zeplin2 has <0.06 c/kev/kg/yr at 2.5MeV

(limit extracted from 2 days realtime of monitoring runs)

800kg detector:

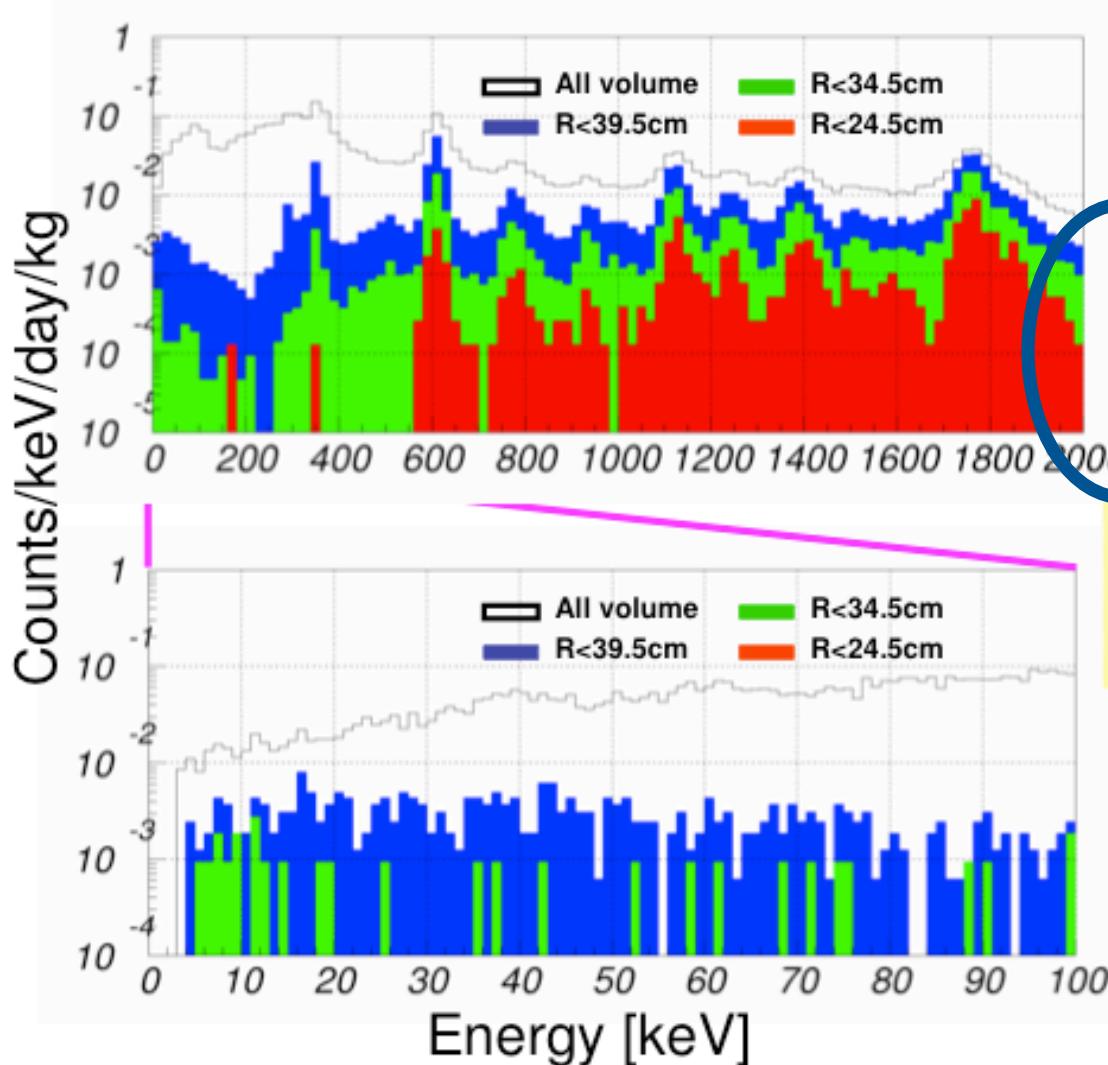
100kg fiducial mass for  $\beta\beta$ -search

Extrapolating  $\Delta E$  from prototype and photocathode coverage:  $\Delta E/E \sim 0.05$

To be competitive: reduction by a factor 10e5  
to  $\sim 1 \text{e-}4 \text{ c/kev/kg/yr}$

# Background estimate for DM search !!

## ➤ Estimation of $\gamma$ ray BG from PMTs



- U-chain
- 1/10 lower BG PMT than R8778

Statistics: 2.1 days

No event is found below 100keV after fiducial cut ( $R < 24.5\text{cm}$ )

$< 1 \times 10^{-4} \text{ cpd/kg/keV}$   
can be achieved  
(Now, more statistics  
is accumulating)

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# EXO

## *Two detector options under consideration*

### High Pressure gas TPC

- 20 atm, 35 m<sup>3</sup> modules,  
4.2 ton/module, 2 modules
- Xe enclosed in a non-structural  
bag
- β range ~5cm: can resolve  
2 blobs
- 2.5m e-drift at ~250kV
- Readout Xe scintillation with  
WLSB (TO)
- Additive gas: quenching and  
 $Ba^{++} \rightarrow Ba^+$  neutralization
- Steer lasers or drift Ba-ion to  
detection region

### Liquid Xe chamber

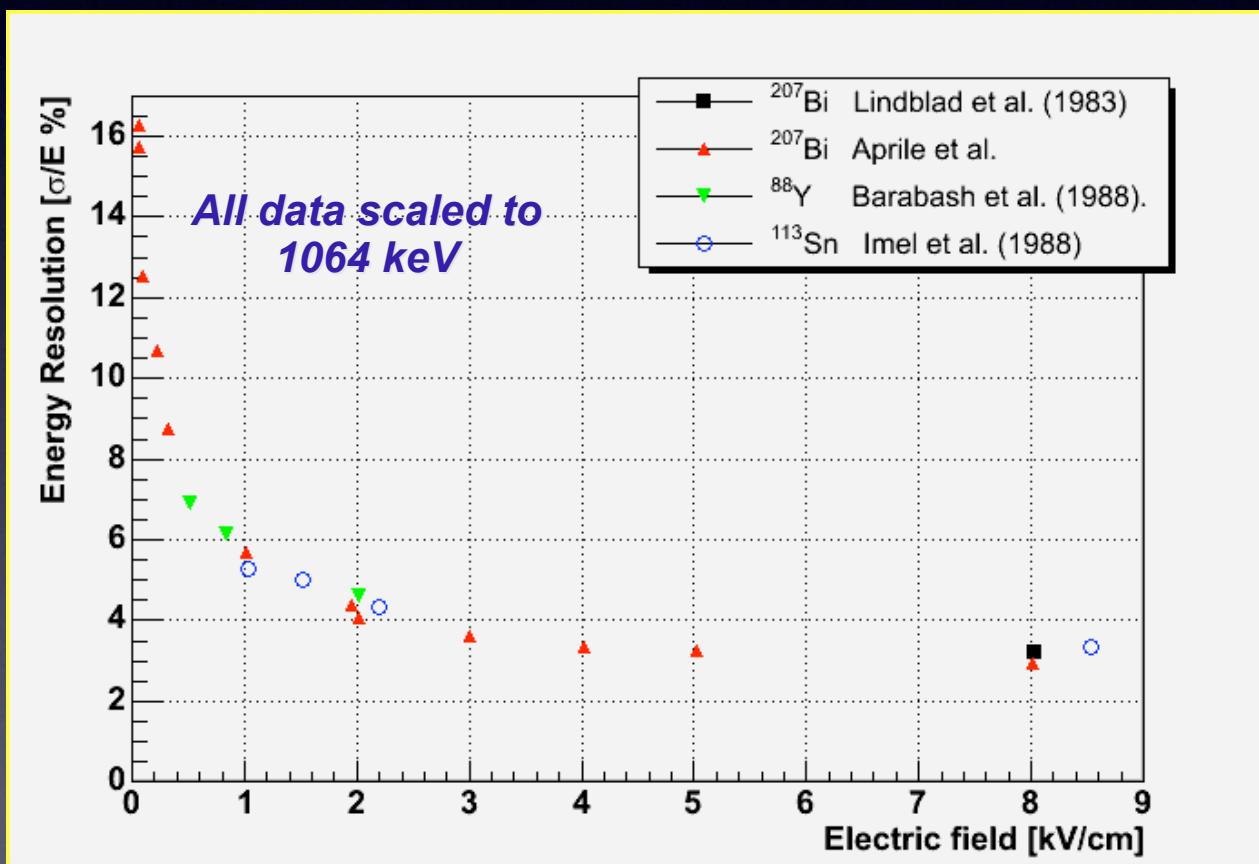
- Very small detector (3m<sup>3</sup> for  
10tons)
- Need good E resolution
- Position info but blobs not resolved
- Readout Xe scintillation
- Can extract Ba from hi-density Xe
- Spectroscopy at low pressure:  
 $^{136}Ba$  (7.8% nat'l) different  
signature from natural Ba  
(71.7%  $^{138}Ba$ )
- No quencher needed, neutralization  
done outside the Xe

# EXO

- Main problems to address:
  - Energy resolution - not least to disentangle 2-neutrinos from 0-neutrino mode
  - Background rejection -  $\beta\beta$  signature

# EXO- energy resolution

Highly non-Poisson statistics limit  
calorimetric resolution in LXe



$$\sigma \sim \sqrt{FN_e}$$

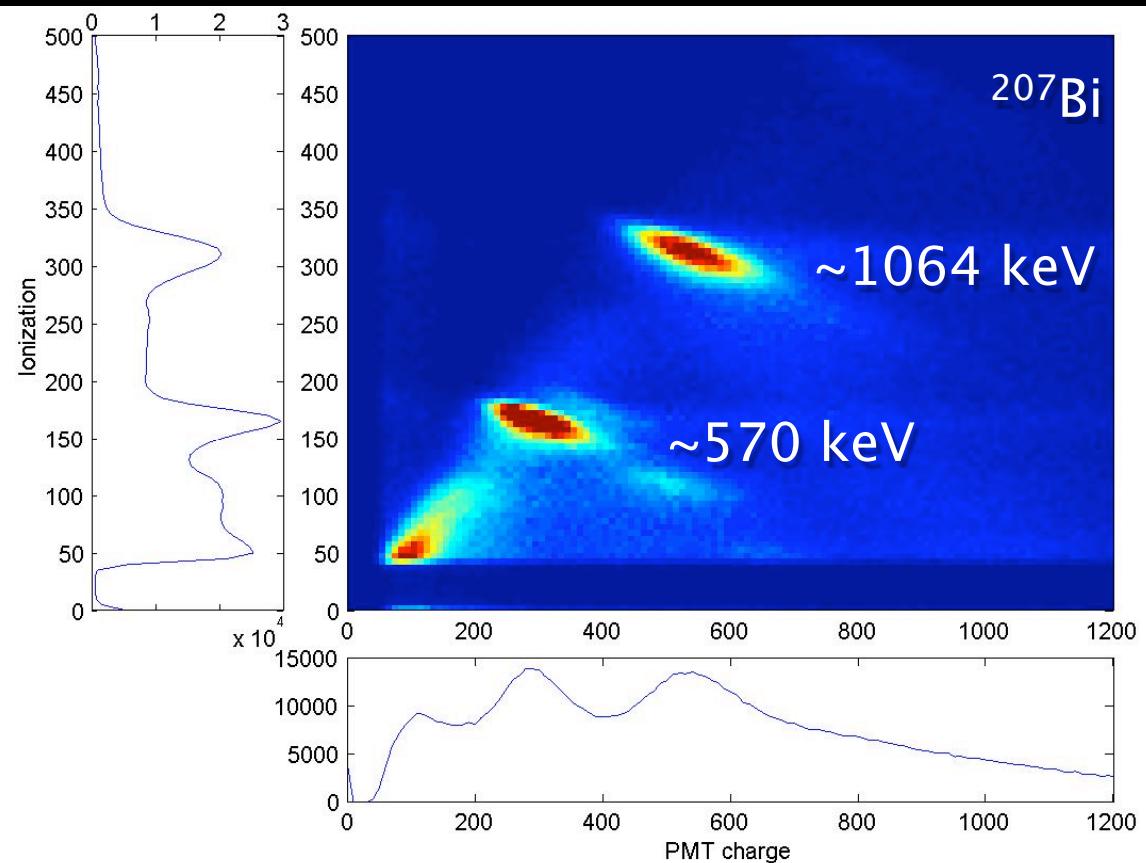
$$F_{\text{Pred.}} \sim 0.05^a$$

$$F_{\text{Meas.}} > 20^b$$

<sup>a</sup> T. Doke, et. al, Nucl. Instrum. Methods Phys. Res. 134, 353 (1976)

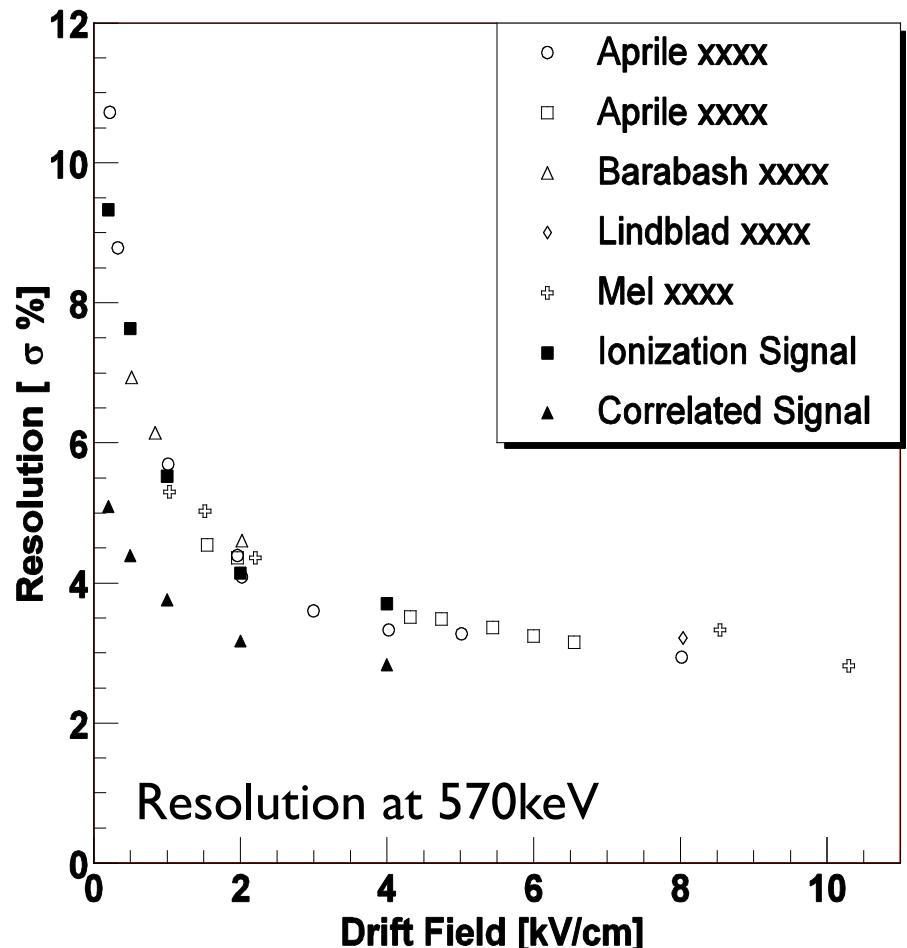
<sup>b</sup> E. Conti et. al, Phys. Rev. B 68, 054201 (2003)

# EXO- energy resolution



Using anti-correlation:

$\Delta E/E \sim 0.014 @ 2.48\text{MeV}$   
(0.018 with ionisation only)



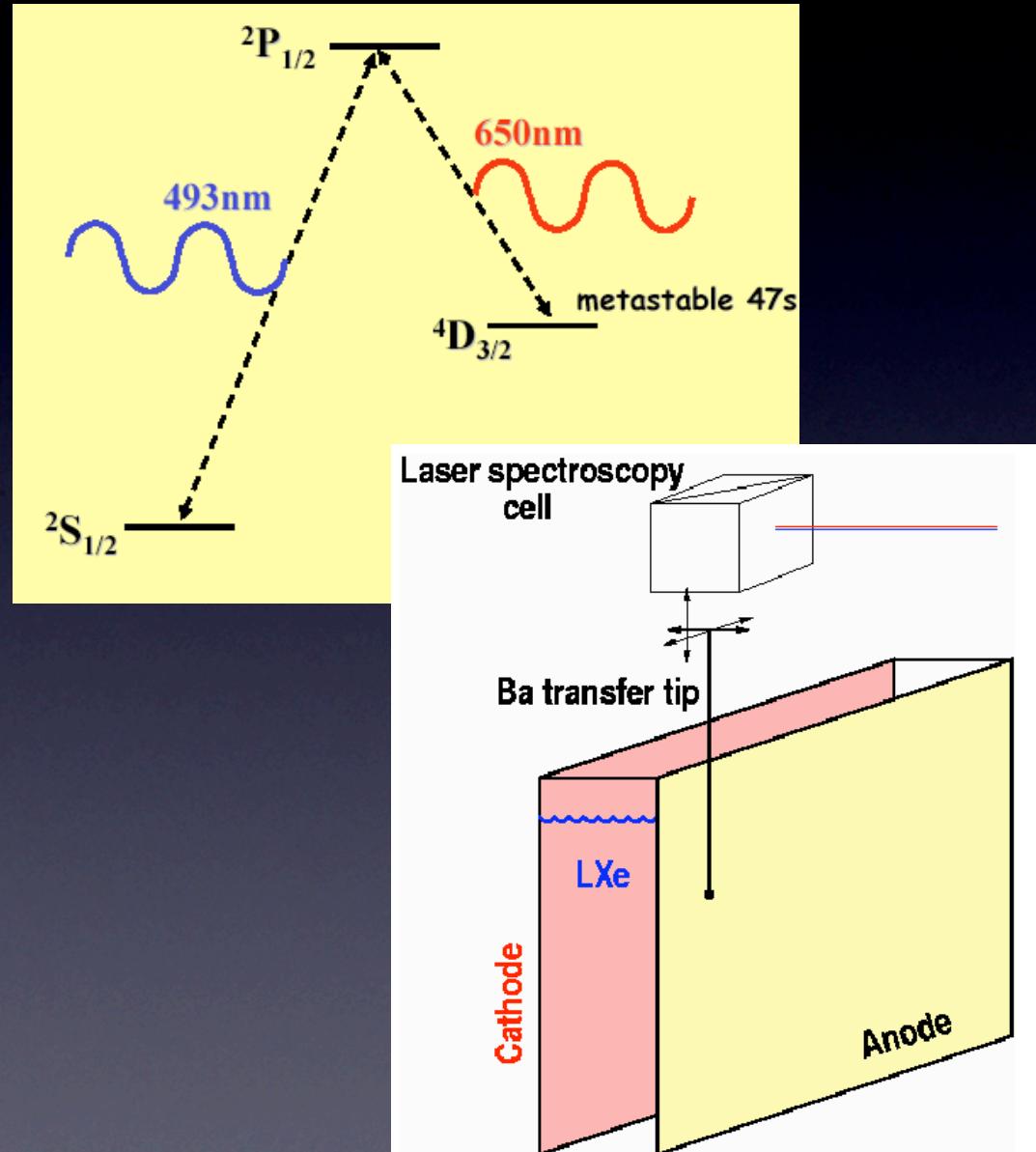
Small test chamber results  
(EXO-200 will collect 3 times more light  
- ie further improvement possible)

# EXO $\beta\beta$ signature

Xe offers a qualitatively

new tool against  
background:

$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}^{++} e^- e^-$  final  
state can be identified  
using optical  
spectroscopy (M. Moe, Phys. Rev.  
C 44 (1991) 931)

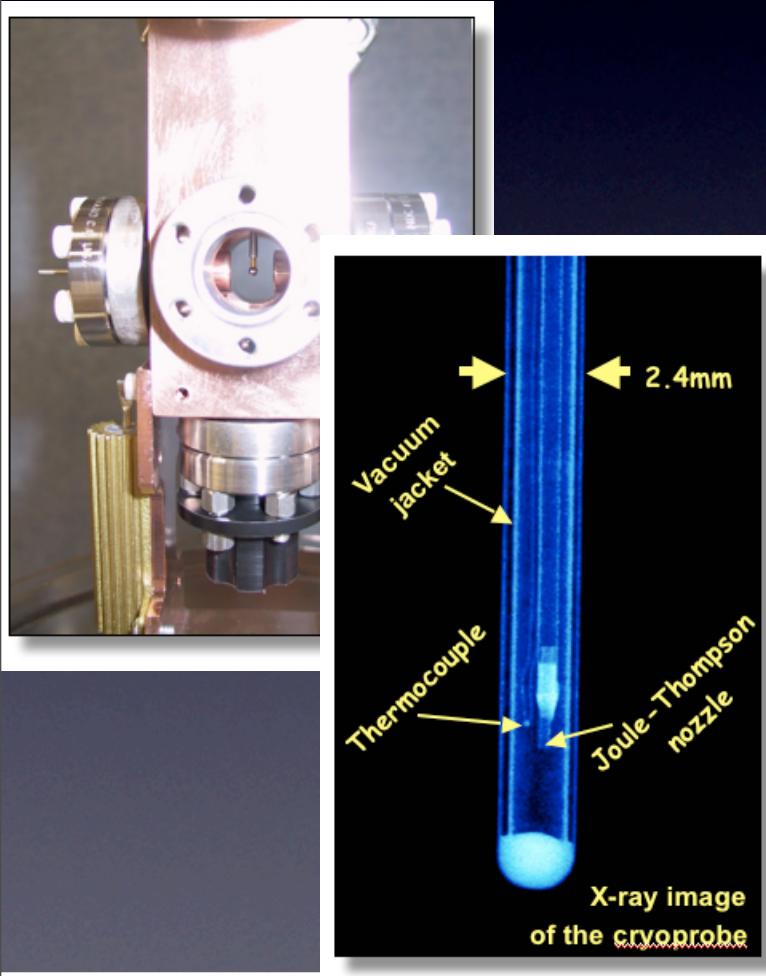


GXe: laser brought to event

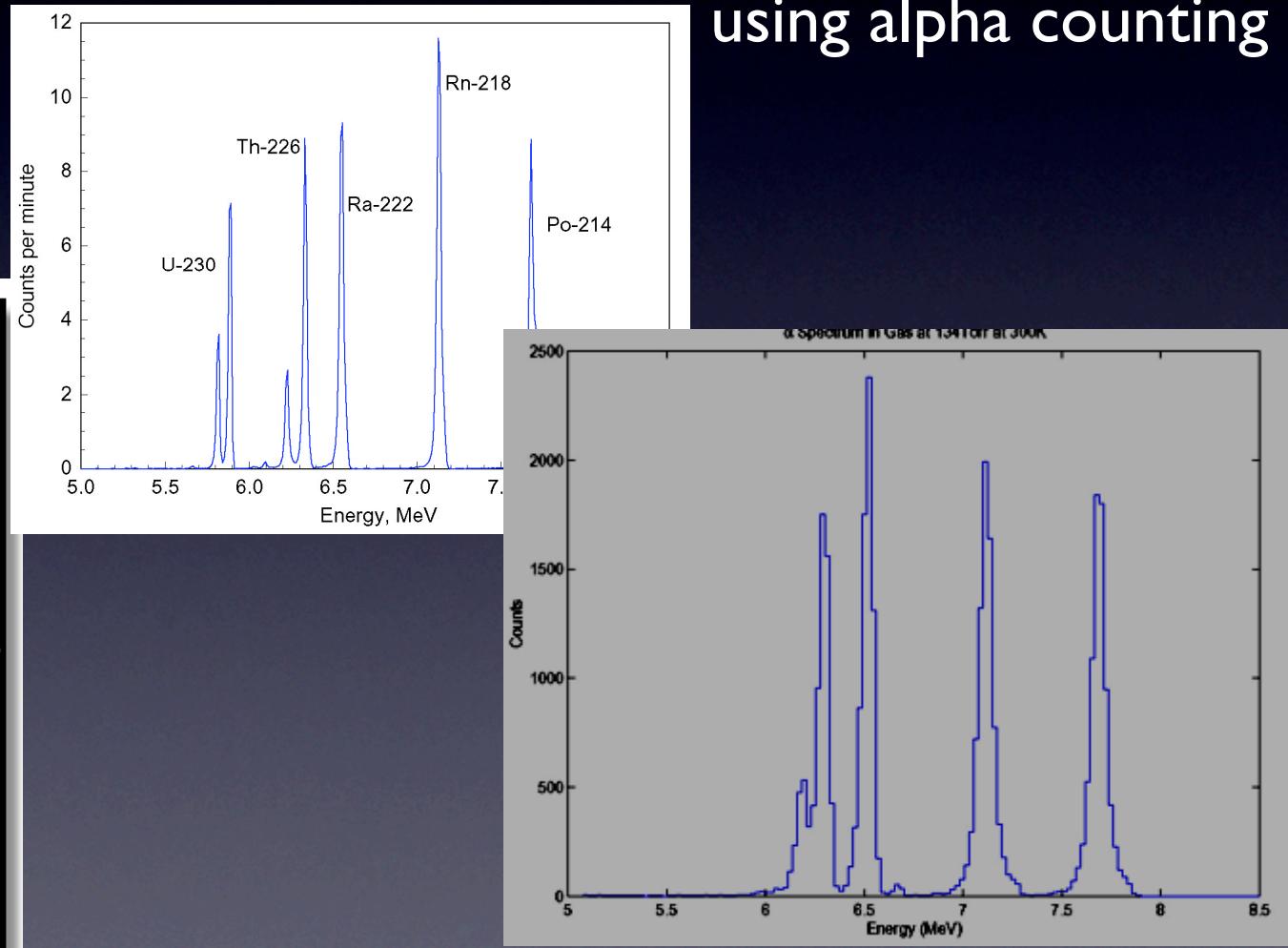
LXe: ion extracted, moved to  
quadrupole and analysed  
Important additional constraint

# Laser tagging - how it works

## I. grabbing of $^{136}\text{Ba}^+$ ion

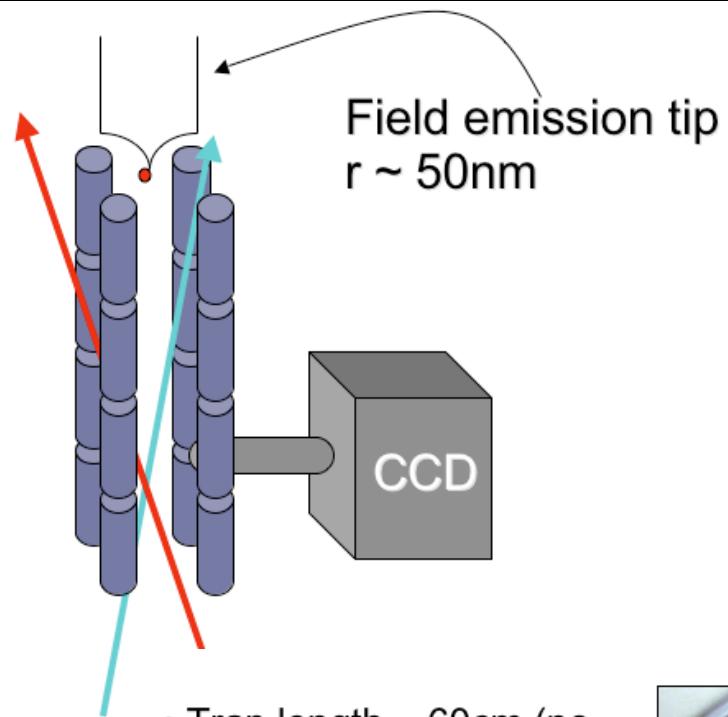


Tested with  $^{230}\text{U}$  ions:  
similar chemistry, but easier signature  
using alpha counting

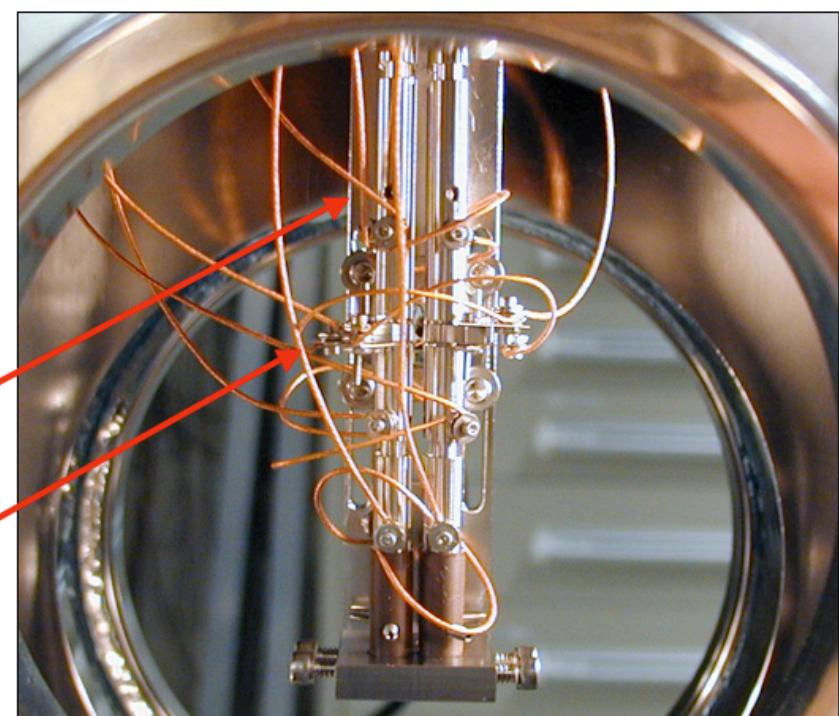


## 2. release of $\text{^{136}\text{Ba}^+}$ in quadrupole

- Load ion into trap with oven
- Observe in trapping region
- Field **adsorb** ion onto tip
- Field **desorb** ion from tip into trap and observe

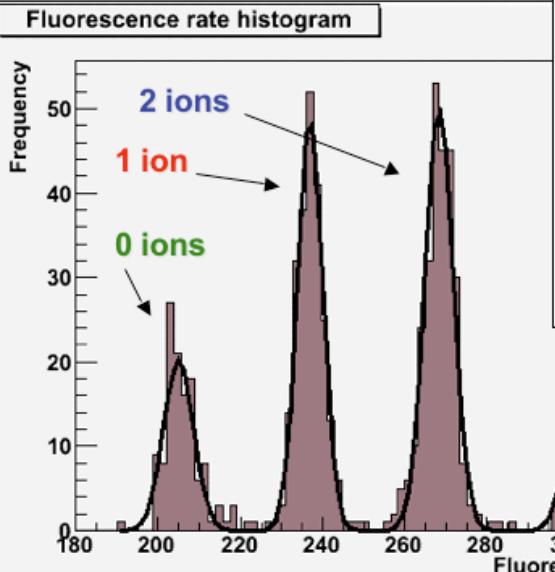


- Trap length  $\sim 60\text{cm}$  (no limiting factor on length)
  - 16 independently computer-controlled segments (real time)
  - RF low (no breakdown in buffer gas)
- Transport segments ( $\sim 3\text{ cm length}$ )
- Trapping segments ( $\sim 3\text{ mm length}$ )



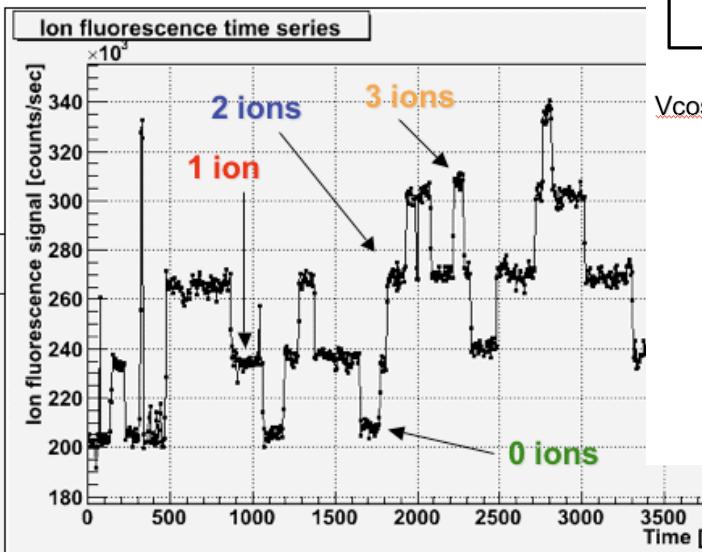
# Ion quantization in Helium

- $P_{\text{buffer}} \sim 10^{-3}$  torr Helium
- 5 sec integration per point
- Peak separation  $\sim 8.7\sigma$



## 3. evidence of $^{136}\text{Ba}^+$ ion

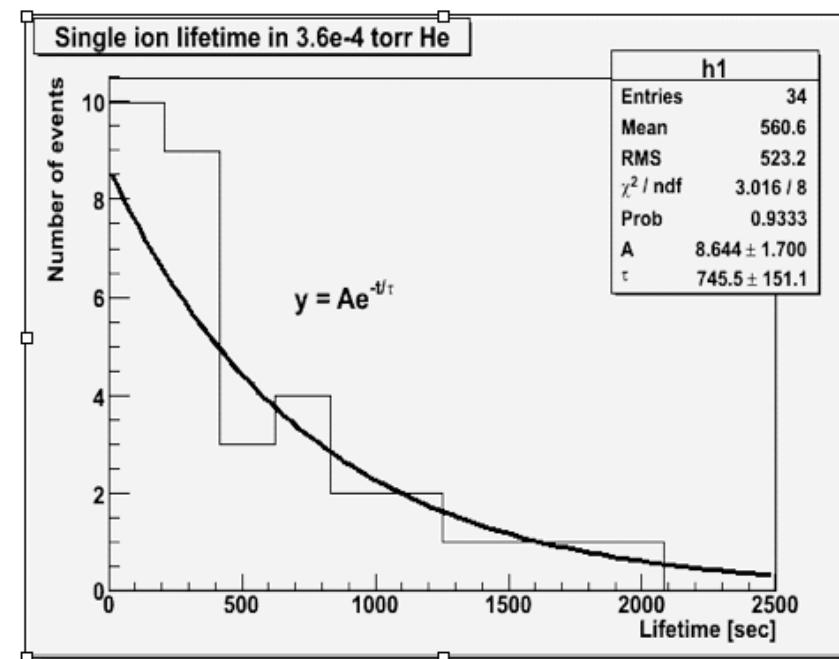
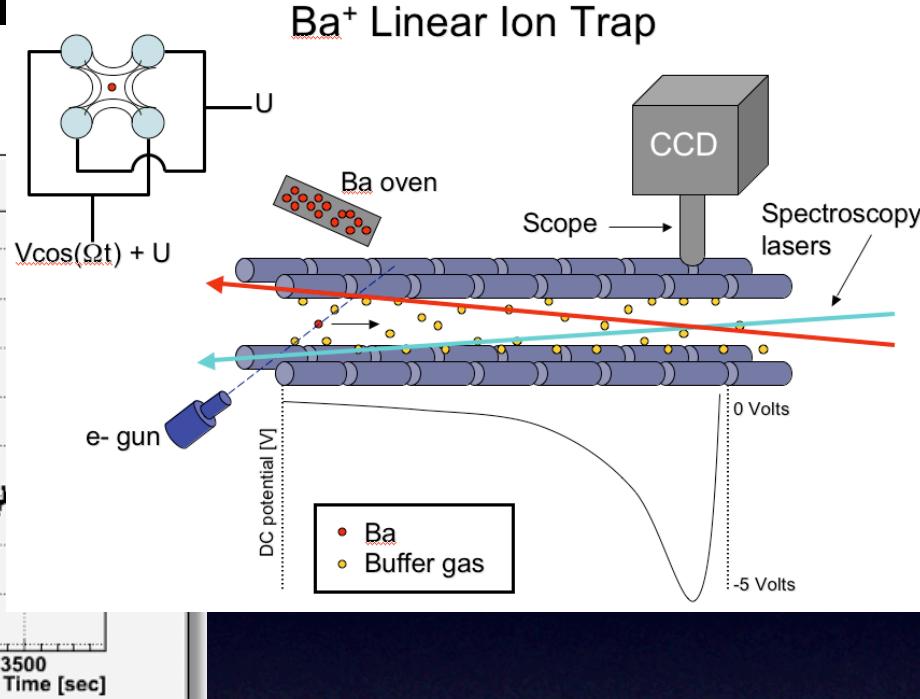
been tested in He and Argon



- $P_{\text{buffer}} \sim 3.6 \times 10^{-4}$  torr He
- Single ion is loaded, and timed until ejection from trap
- Ion lifetimes histogrammed and fit to exponential
- Lifetime follow exponential distribution with

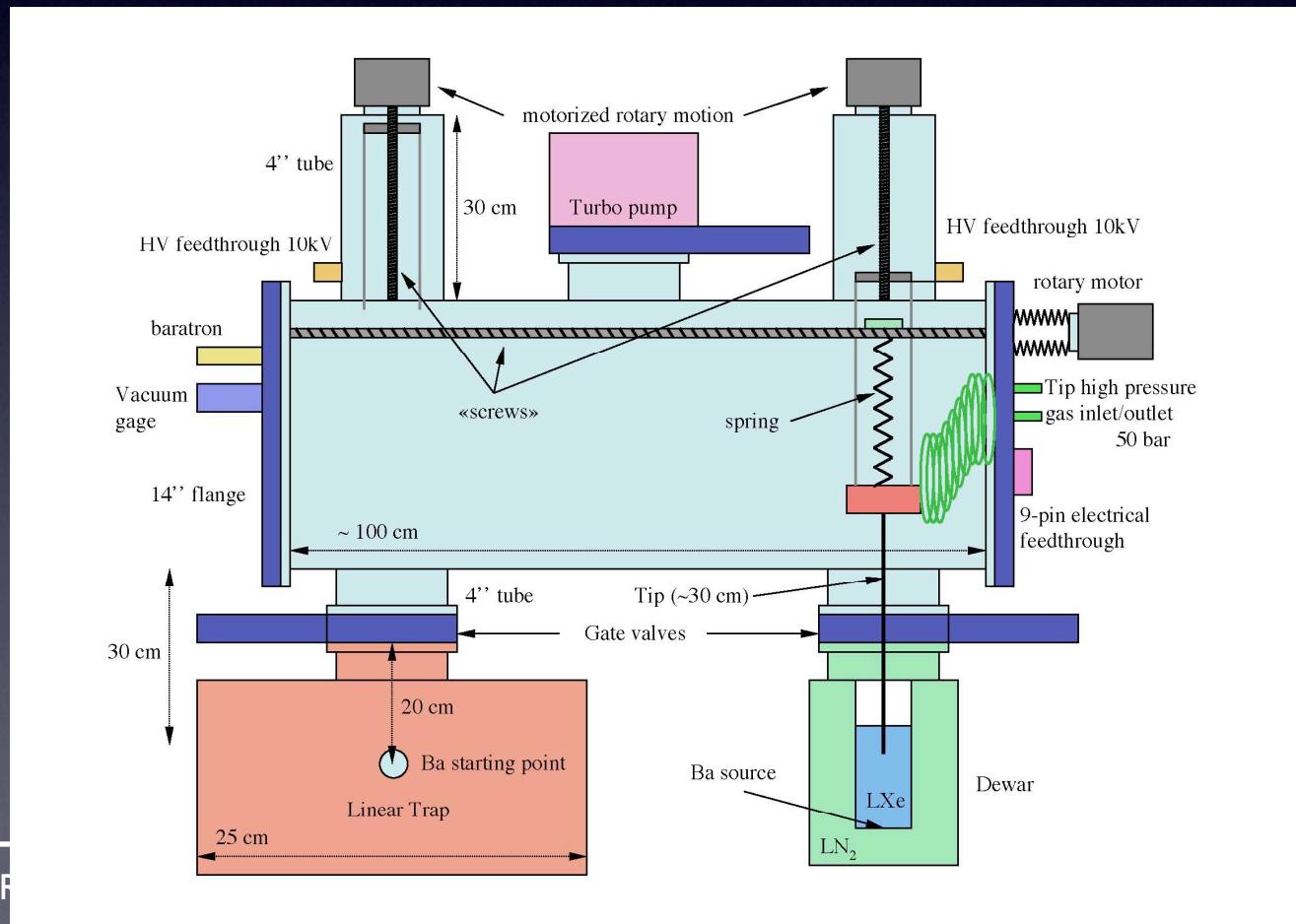
$$\tau \sim 746 \pm 151 \text{ sec}$$

→ Probably capture on impurities in trap ( $\text{O}_2$ ,  $\text{NO}$ ,  $\text{CO}_2$ , etc.)



# Laser tagging - R&D

- Choice of grabbing tip
  - QCM tip; FE tip (field emission); RIS tip (resonant ionisation spectroscopy)
- Proof of grabbing/releasing/measure all-in-one in a ‘real’ detector
  - build of an R&D 100kg LXe TPC



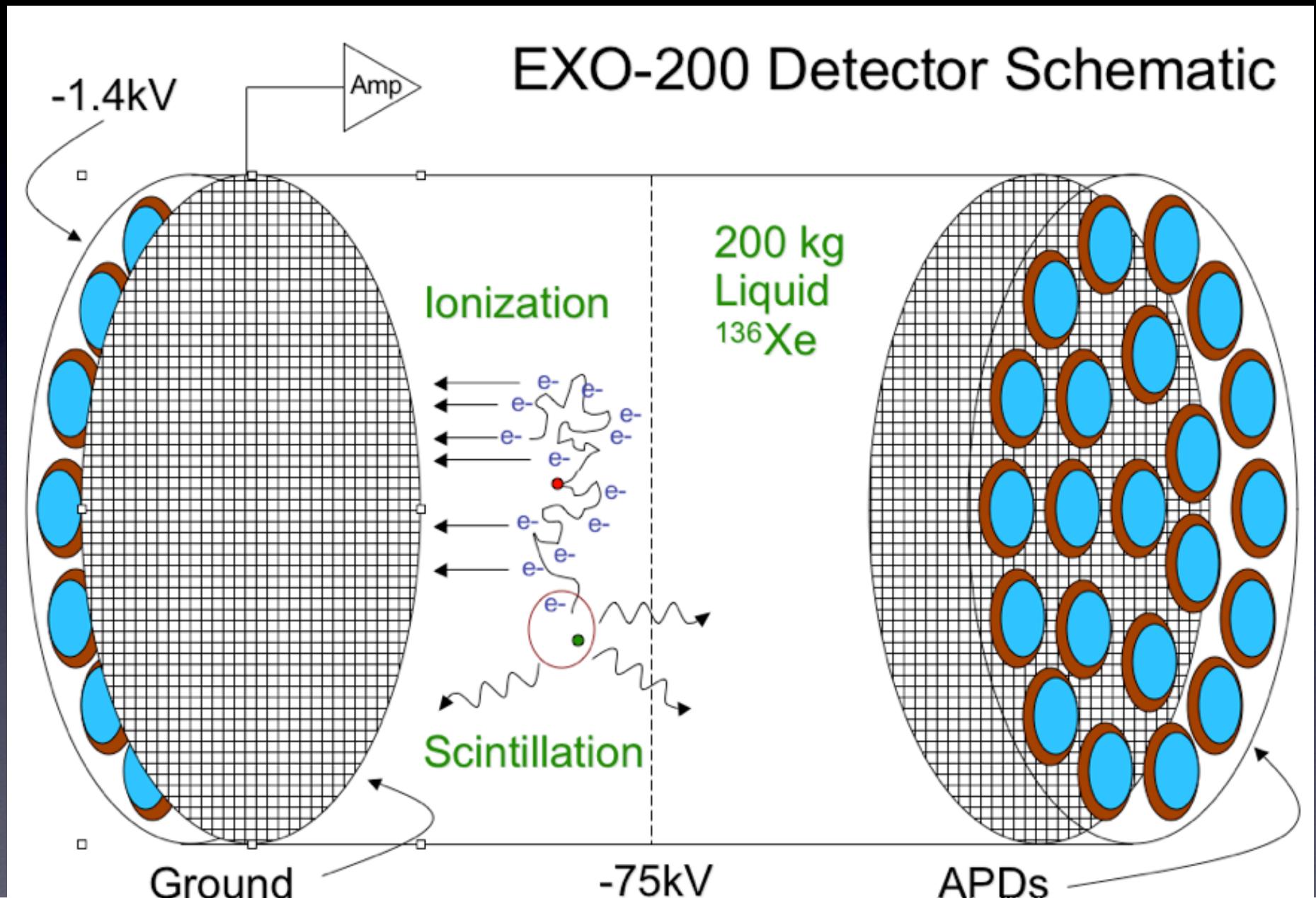
# How EXO works

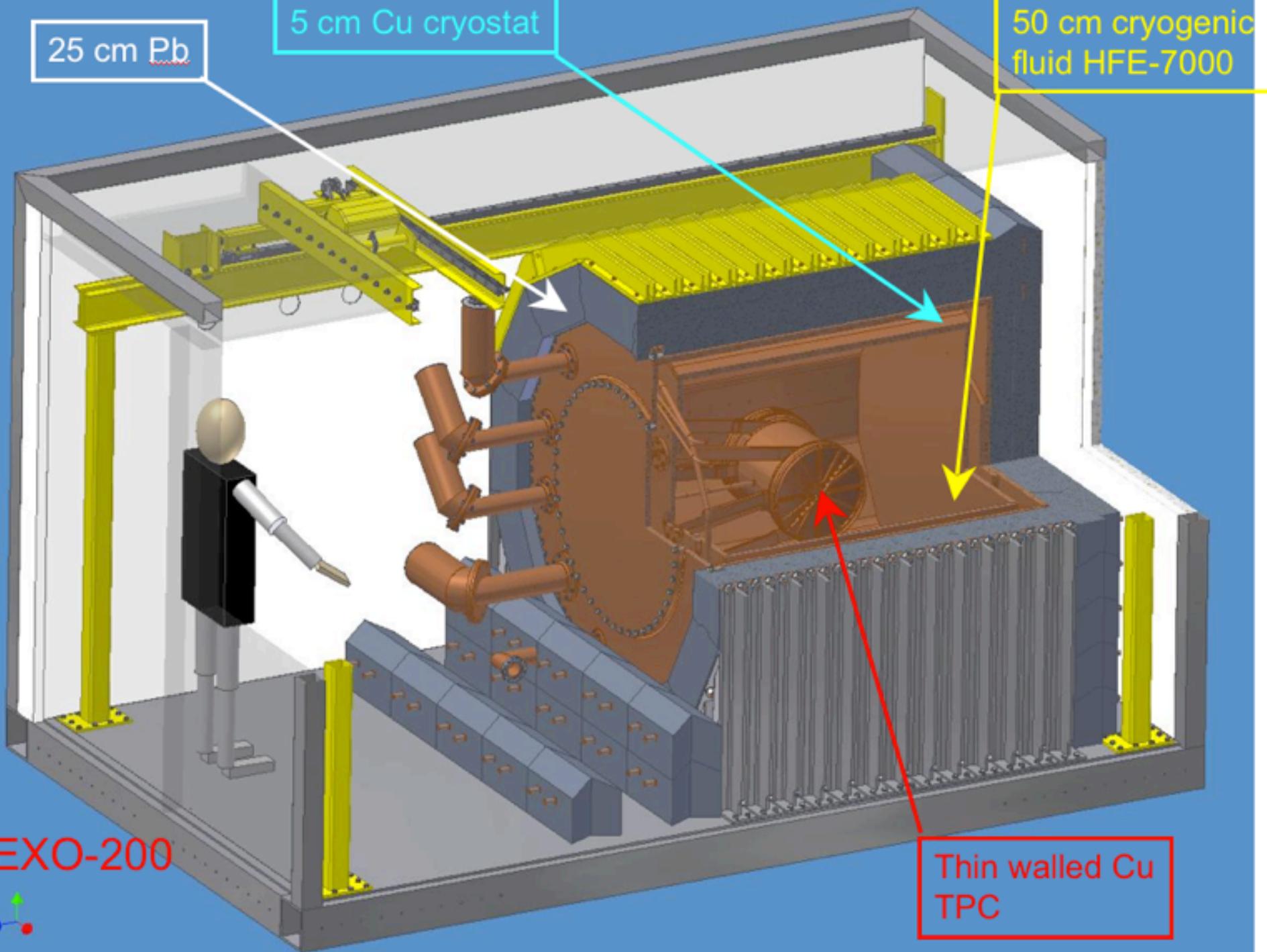
- Isotopically enriched liquid  $^{136}\text{Xe}$  (80%)
  - Semi-cryogenic (-115 °C)
  - LXe acts as both source and detector
- $2\nu\beta\beta$ ,  $0\nu\beta\beta$  (and background events) create ionization trails in LXe ( $e^-$  and  $Xe^+$ )
  - $W_{Xe} \sim 10 \text{ eV}$ ,  $Q_{\beta\beta} \sim 2.5 \text{ MeV} \rightarrow \sim 10^5 e^-/\text{event}$
- Drift field sweeps up most electrons (TPC)
  - Sum up  $e^-$  (no gain in liquid) gives energy of decay
  - Pulse height/timing analysis gives x-y position of decay (Res.  $\sim 1\text{cm}^3$ )
- Some  $e^-$  and  $Xe^+$  left over recombine and scintillate (175nm)
  - Timing gives a “start,” and allows for z-position reconstruction
- $\text{Ba}^{++} \rightarrow \text{Ba}^+$  left over drifts in opposite direction

EXO-200

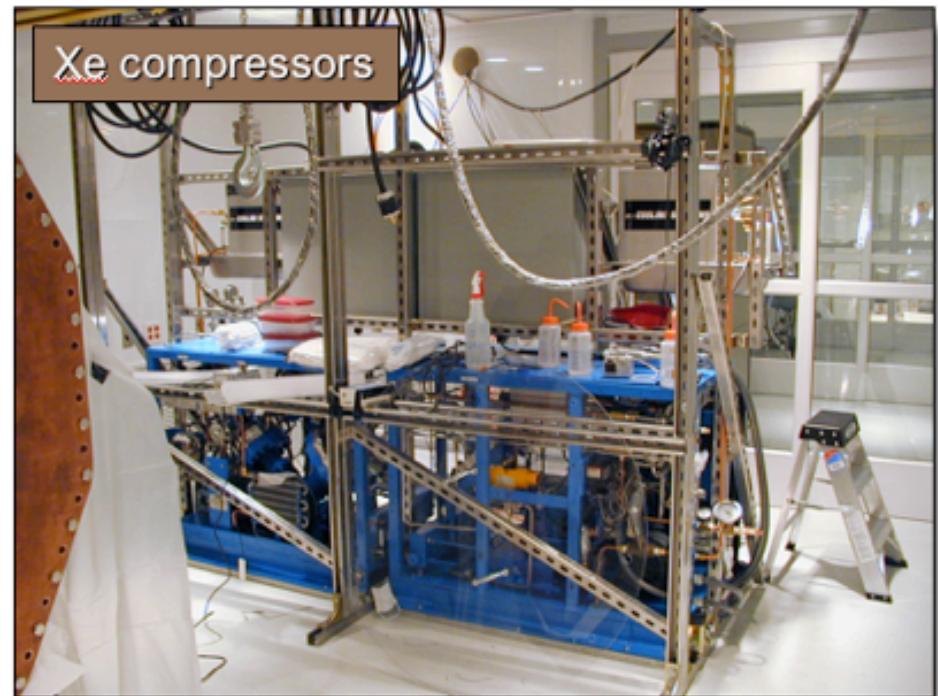
Full EXO

# EXO-200 - LXe TPC

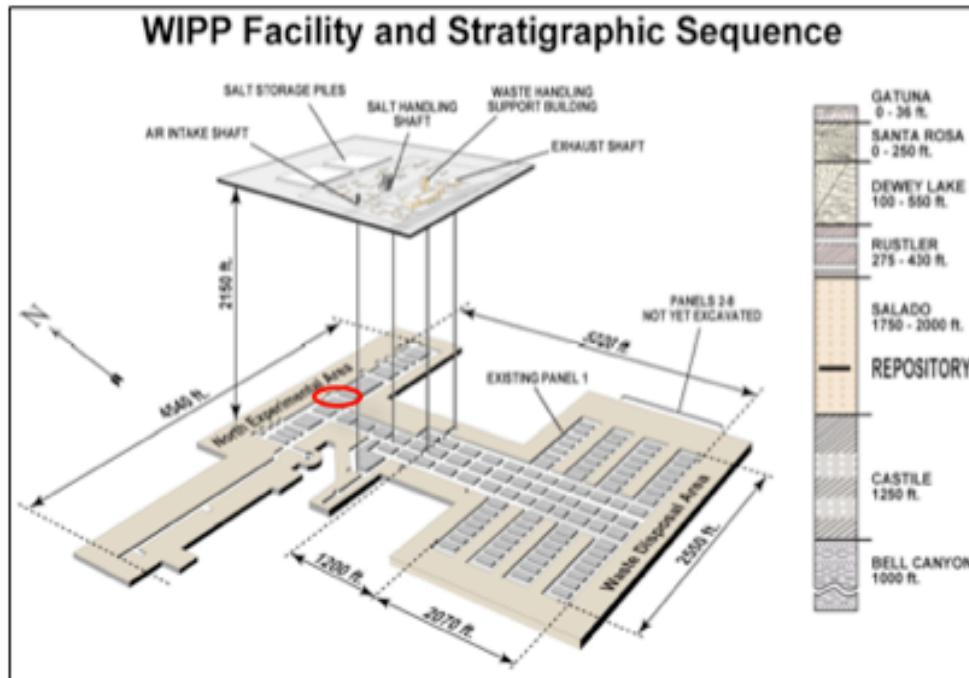




# EXO-200 commissioning at HEPL, Stanford



# EXO-200 Experimental Site



Installation ancillaries started  
Installation of detector March 07



- 655 m underground (rock + salt), 1600 m.w.e
- vertical muon flux =  $3.1 \times 10^{-7} \text{ s}^{-1} \text{ cm}^{-2} \text{ sr}^{-1}$

(NIMA 538 (2005) 516)

# EXO - sensitivity

Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma_E/E$ @ 2.5MeV (%)	Radioactive Background (events)	$T_{1/2}^{0\nu\beta\beta}$	Majorana mass (eV)	
						(yr, 90%CL)	QRPA	NSM
EXO-200	0.2	70	2	1.6*	40	$6.4 \times 10^{25}$	0.27 <sup>†</sup>	0.38 <sup>♦</sup>

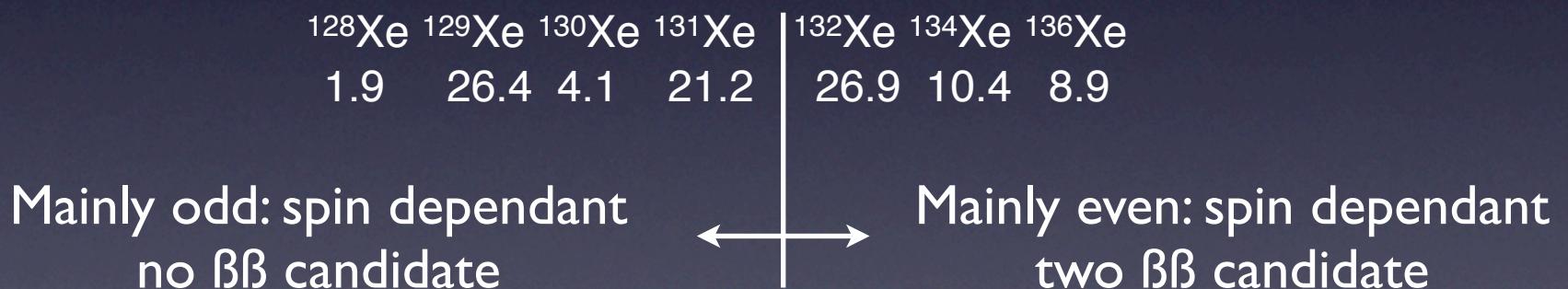
## Future - with laser tagging

Case	Mass (ton)	Eff. (%)	Run Time (yr)	$\sigma_E/E$ @ 2.5MeV (%)	$2\nu\beta\beta$	$T_{1/2}^{0\nu}$	Majorana mass (meV)	
					Background (events)	(yr, 90%CL)	QRPA <sup>‡</sup>	NSM <sup>#</sup>
Conservative	1	70	5	1.6*	0.5 (use 1)	$2^*10^{27}$	50	68
Aggressive	10	70	10	1 <sup>†</sup>	0.7 (use 1)	$4.1^*10^{28}$	11	15

# Combining DM & $\beta\beta$ ?

... as both fields seek ton-scale experiment?

- Bargain? Buy one get, one free?!?
- No really: cost for enriched Xe, essential for  $\beta\beta$   
= cost for rest of DM experiment
- WIMP spin independent/dependent cross section studies  
**would anyway require some isotopic separation!**



- ‘No waste’ enrichment, ie cheaper